First meeting: Coronal Magnetism ISSI team

February 25 - March 1 2013

Presentations

Available at http://www.issibern.ch/teams/coronalmagnet/firstmeeting.html

VLA/CoMP/Hinode/(SolarC?) observing run in April

• Tim/Sarah/Steve/Alphonse/(Jeff?) - maybe also Terry/Don S/Kathy for analysis depending on Hinode data

Expansions/applications of FORWARD

- DONE! Added green line to FORWARD
 - set instrument='greencomp'
- Add radio to FORWARD

• free/free and gyro - **Tim** will send function to **Sarah**, who will incorporate them in FORWARD tree (see below for more details)

- apply to VLA/CoMP/Hinode/(SolarC) observing run
- other diagnostics (e.g. ray-tracing) Tim/Stephen/Christophe
- Add other permitted EUV lines
 - Silvano and Alessandro will write function as described below
 - Sarah will incorporate in FORWARD
- Add visible lines (e.g. 10830)
 - Nour will write function as described below
 - Sarah will incorporate in FORWARD
- Add and analyze new models/simulations Mei/Sarah/Urszula

• compare PFSS, magnetofrictional, and MHD to CoMP and SolarC data (Lol and direction) Cooper/Laurel/Jeff

flexibility/clarification on which spectral model should be used Terry/Kathy/
Cooper

- FORWARD widgetization (Sarah/Blake should be done by summer)
- FORWARD solarsoft package (Sarah/Blake should be done by fall)
 ideally, replace FORTRAN CoMP code with IDL (Phil/Roberto)

• Instrument personality profiles -- e.g. CoMP/COSMO (Steve) SolarC/DL-NIRSP (Jeff), ...

Coronal Magnetism Advocacy: what we can do

- General goal: grow a broad user base
 - train generation that will use future telescopes
 - · connect to coronal observer user base, as well as spectropolarimetric
- General outreach methods
 - email exploder
 - · bibliography attached to ISSI web page
- FORWARD analysis tools
 - target multiple user levels (including "gateway")
 - data products like forward-modeled PFSS/MHD/magnetofrictional
 - REU type project development
 - sit-down/hands-on workshop to train new FORWARD users
- White paper/letter
 - content
 - science objectives -- why we want to know coronal B
 - assets, current and future emphasize complementarities
 ATST, COSMO, CSRH, FASR, SIGMA, SOLARC
 - how we would put these together to obtain coronal B
 - FORWARD (tomography? forward-fitting?)
 - audience
 - NASA, ESA, JAXA
 - NSF (Coronal magnetism program?)
 - SPD, CSSP
 - NSO, HAO/NCAR
 - ISSI directors, COST
 - why us?
 - harmonic convergence of broad set of stake-holders
 - · would be good to get more community input
- · Meetings, workshops upcoming opportunities
 - AGU May 2013 (poster on FORWARD)
 - IAU prominence meeting June 2013 (presentations)
 - SHINE June 2013 (session?)
 - SPD July 2013 (?)
 - SPIE August 2013 (?)
 - SPW September 2013 (presentations?)
 - AGU December 2013 (session?)
 - Hinode November 2013 (?)
 - ISSI workshop/textbook -- propose after next team meeting
 - · COST/HAO Coronal Magnetism workshop series?

To add an observable, you need to create an IDL function as follows (example for emissivity):

FUNCTION to calculate emissivity along the LOS

INPUTS TO FUNCTION (provided by FORWARD - generally two-dimensional, LOS vs POS)

r3d theta3d phi3d	
	(spherical coordinates, central meridian at phi = 0)
Brobs	
Bthetaobs	
Bphiobs	
Velr	
Veltheta	
Velphi	
Dens	
Temp	

FUNCTION SHOULD RETURN

emissivity (three-dimensional, LOS vs POS vs 4 (Stokes vector))

Comments:

*Radio works very much as emissivity -- but, my expectation will be that Radio will send back two-dimensional arrays (LOS vs POS) having done the integral, but emissivity will not.

*In order to test your function, you can use a datacube generated e.g. as follows (verbose forces the dump of data):

For emissivity off-limb:

for_drive,'giblow',instrument='euvia',line='195',xxmin=0.8,xxmax=1.6,yymin=. 5,yymax=.5,losuse='XLos',out=1,memory=0,ngrid=64,/verbose

For radio -- including disk:

for_drive,'giblow',instrument='euvia',line='195',xxmin=0.8,xxmax=1.6,yymin=. 5,yymax=.5,losuse='XLos',out=1,memory=0,ngrid=64,/verbose

This will generate an IDL save file called "datadump" with the INPUT variables.