

Spectral lines: synthesis

- Fe I 6301+6302+6303, Fe I 5247+5250, Fe I 15648+15652, Fe I 6173, Mn I 5537, Mn I 1.52, Si I 10827
- SPINOR (Michiel, Andreas), SIR (Reza, JM), LILIA-NICOLE (Nikola, Hector), MISMA (Bartolomeo)
- Andres gathers and averages

Spectral lines: inversions full MHD resolution

- Employing simulations: Arturo vs Thorsten:
ALL LINES different snapshot for training and inversion
- Tau-dependent: Ex 1) Fe I 6301+6302; Ex 2) Fe I 6301+6302+6303, Fe I 5247+5250, Fe I 15648+15652
- ME: Ex 1) 6302; Ex 2) Fe I 6301+6302
- Spectral range for 6301+6302: 6301-6303

Code users

- SPINOR (Michiel, Andreas), SIR (Reza, JM), LILIA-NICOLE (Nikola, Hector), MISMA (Bartolomeo)
- HAO/ASP (Bruce), VFISV (JM), Helix++ (Andreas)
- Arturo (tbd) and Thorsten (IMAP)

Inversion full MHD resolution

- 1 component + no filling factor
- MISMA possible ?
- Freedom of selection for nodes positions and number, and weights
- Same for other fudging parameters: macro, micro, damping enhancement
- Same atomic parameters and abundances: JM provides
- We all save the best-fit profiles
- Same wavelength ranges (also ME codes among themselves)

Non ideal examples: Instrument case Spectrograph

- 0.3"; 0.1"; 1" (QS)
- 1E-3; 1E-4
- QS+Plage+Sunspot
- 6301+6302
- Include by default a non-magnetized atmosphere
- Total: 14 inversions

Non ideal case Instrument filter

- 0.1", 0.3"
- QS+Plage+Sunpot
- 1E-3
- 5250 (5 and 12 points; 65 mA FWHM with 60 and 32 mA sampling, respectively)
- Total: 12 inversions

- 1 min pixel
- 50000 pixels = 50000 min = 833 hours = 35 days (no crash, 1 cpu)
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Priorities # 1

- 3000*16 column from Matthias: 16*16*12 km
- Synthesis
- Inversion full resolution
- All codes
- QS: 1536*1536 at 24*24*12 with mixed polarity: +10, -10.

Priority # 2

- For those codes that are able to: use ME and tau-dependent inversions for a 2 magnetic component inversion in some selected regions of wierd profiles (e.g. Penumbra)
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Further test for ME before inversion

- Produce a few more cases to make sure codes give the same numbers. Compare also shape of profiles
- $B=10, 100, 500, 1500, 3000$
- $G=0, 45, 90$
- $\Phi=30$
- $V_{los}=0$
- 10 mA, -500, +1500, 6301.5D0, 6302.5D0
- $A=0.1$
- $DLDOP=30$
- $ETA0=10$ (for 6302.5); ratio=2.80D0
- $S0=S1=0.5$

Further tests for tau-dependent codes before inversion

- Opacity tables with $T=2500$ to 10000 in 500 K steps
- $\text{Pelec}=1\text{E}-3$ to $1\text{E}3$ in decades in log steps
- $\text{Lambda}= 5000, 15000$

Comparison with simulations

- For tau-dependent: $\log\tau=0,-1,-2,-3,-4$. Ask Thorsten and Arturo to convert to Tau.
- For ME: use RF to temperature ? Calculate them while synthesizing with SIR
- For spatially averaged simulations do not do anything. Will be decided later.

Schedule

11 February: web to exchange data

- MHD cubes: $3000 \times 16 \times 192$
- Presentations
- Atomic data and wavelength ranges
- Schedule
- Group picture
- Nikola's paper
- 11 March: synthesis for full resolution [$\lambda, 4, 3000, 16$], split in regions, I/Q/U/V
 - Include 5 wavelength files with exact value used for synthesis
 - Input: FITS
- + 2 week approx: Andres averages
 - Report
 - Output: same format (FITS) and dimensions
- Deadline for inversion results Experiment 1: mid May => results to JM