

H_2 formation and PDR model

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H_2 formation in PDR model

$$dn(H_2)/dt = \alpha n(H) n_H$$

- “Jura 1974“ formation rate $\alpha = 3 \cdot 10^{-17} \text{ cm}^3 \text{s}^{-1}$
- Le Bourlot et al. 95 : $\alpha = 0.5 \langle n_g \sigma \rangle v_H$ calculated for a Mathis power law distribution of grain sizes

$$\langle n_g \sigma \rangle = \frac{\frac{3}{4} \cdot \frac{1.4m_H G}{\rho_g} \cdot \frac{1}{\sqrt{a_{min} a_{max}}}}{= 1.95 \cdot 10^{-21} \text{ cm}^2, v_H = 1.45 \cdot 10^4 T^{0.5} \text{ cm/s}}$$

with $G = 1/100$, $\rho = 3 \text{ g cm}^{-3}$, $a_{min} = 0.03 \mu$, $a_{max} = 0.3 \mu$, $\alpha = 1.4 \cdot 10^{-17} T^{0.5} \text{ cm}^3 \text{s}^{-1}$

- A_v dependent formation rate from Herma studies

$$\eta = a_0 \exp(a_1 \tanh(a_2 A_V))$$

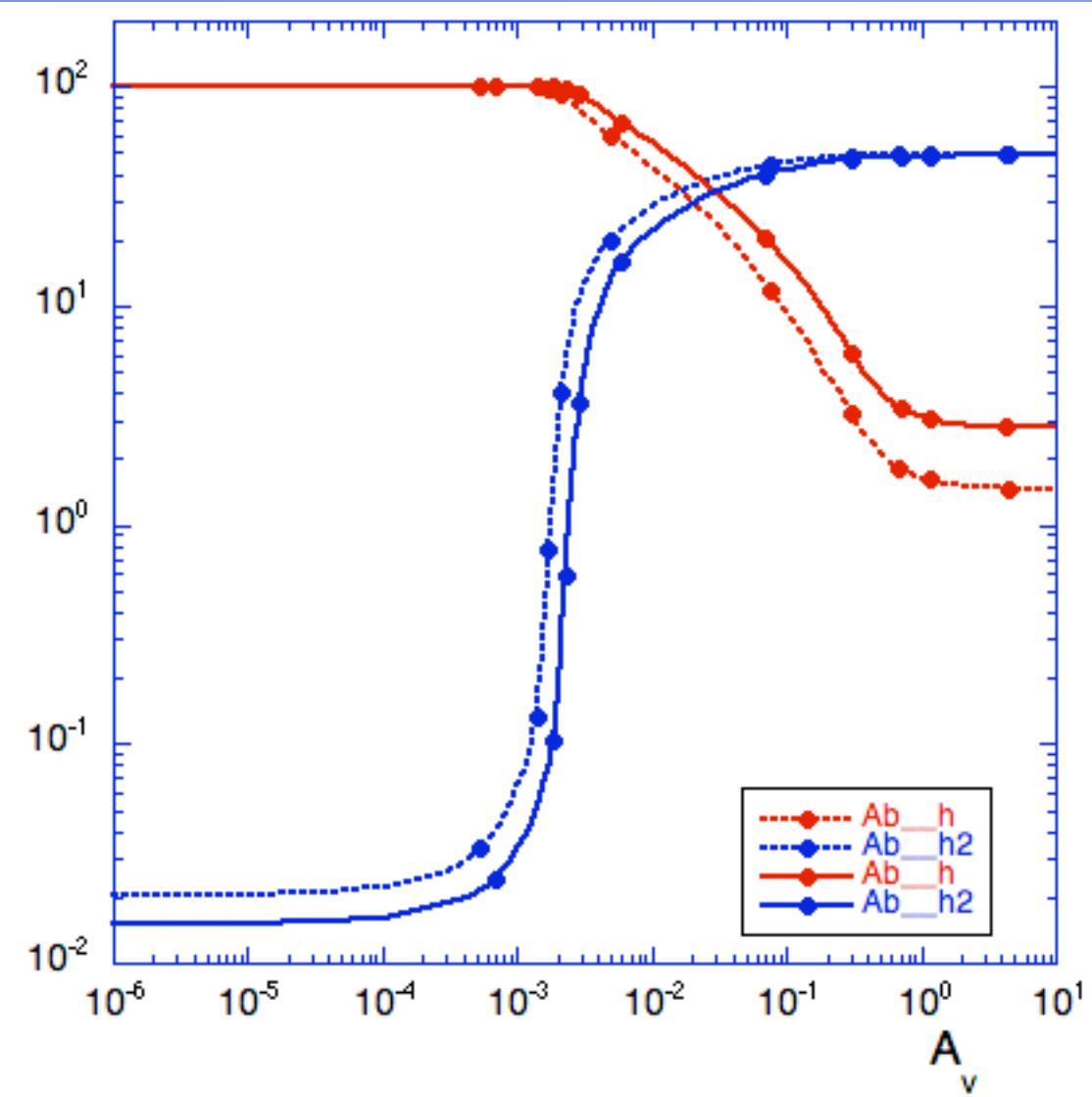
$$a_0 = \exp(\min(2.13 \arctan(281(r - 0.02)) - 3.656, 62000(r - 0.01) - 8))$$

$$a_1 = \min(24.2 \exp(-97.18r), 1423r - 4.11)$$

$$a_2 = \max(-134r + 1.43, 299r - 2.88)$$

Comparison between “standard” and A_v dependent formation rate (I)

- $n_H = 100 \text{ cm}^{-3}$
- $\chi = 1$
- $T = 60\text{K}$
- $\zeta = 5 \cdot 10^{-17} \text{ s}^{-1}$
- Full line : A_v dependent formation rate
- dotted line “standard” formation rate



Comparison between “standard” and A_v dependent formation rate (II)

Molecule	$A_v = 0.7$	$A_v = 1$	$A_v = 2$	$A_v = 3$
H	7.94(19)	8.89(19)	1.18(20)	1.46(20)
	1.27(20)	1.45(20)	2.00(20)	2.53(20)
H_2	6.10(20)	8.84(20)	1.80(21)	2.71(21)
	5.87(20)	8.56(20)	1.76(21)	2.66(21)

1st line: results from Le Bourlot et al. prescription

2nd line : A_v dependent term of H. Cuppen

- Factor of 2 on H column density result
- negligible effect on molecular content
- Extension to dense photodissociation regions such as the Orion bar (role of dust temperature)