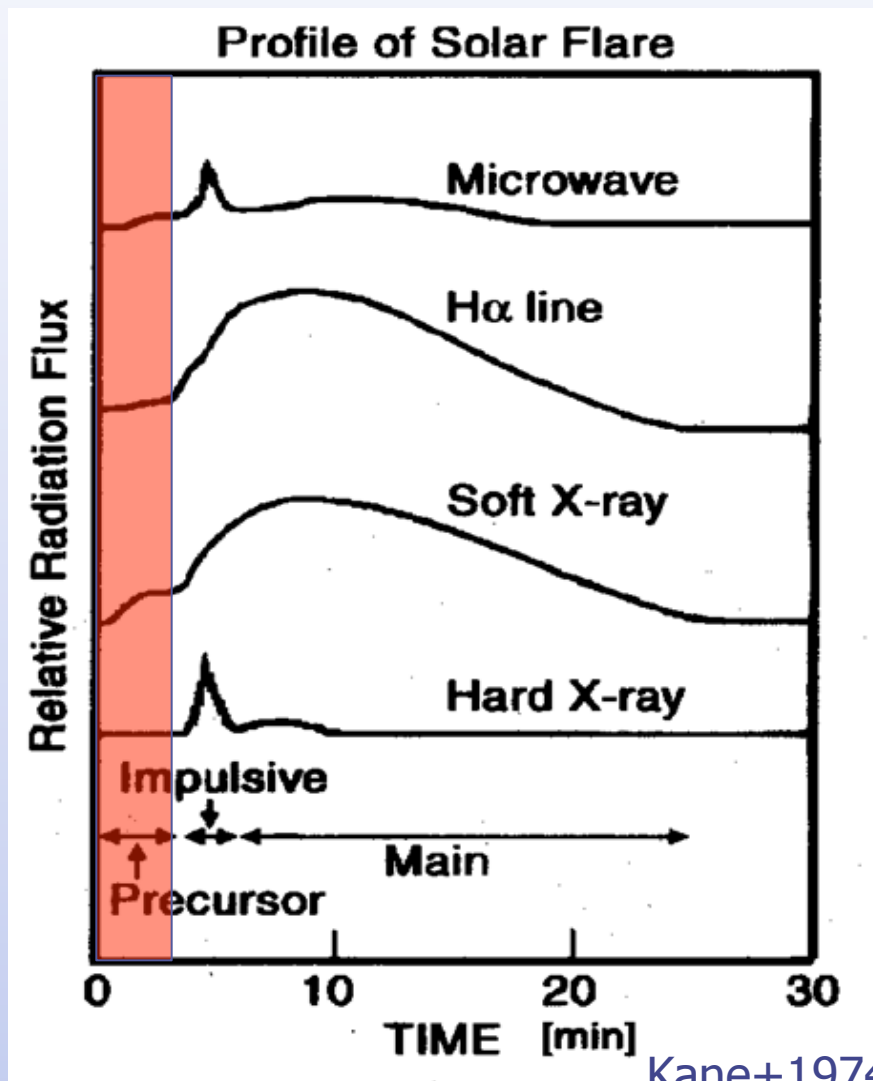
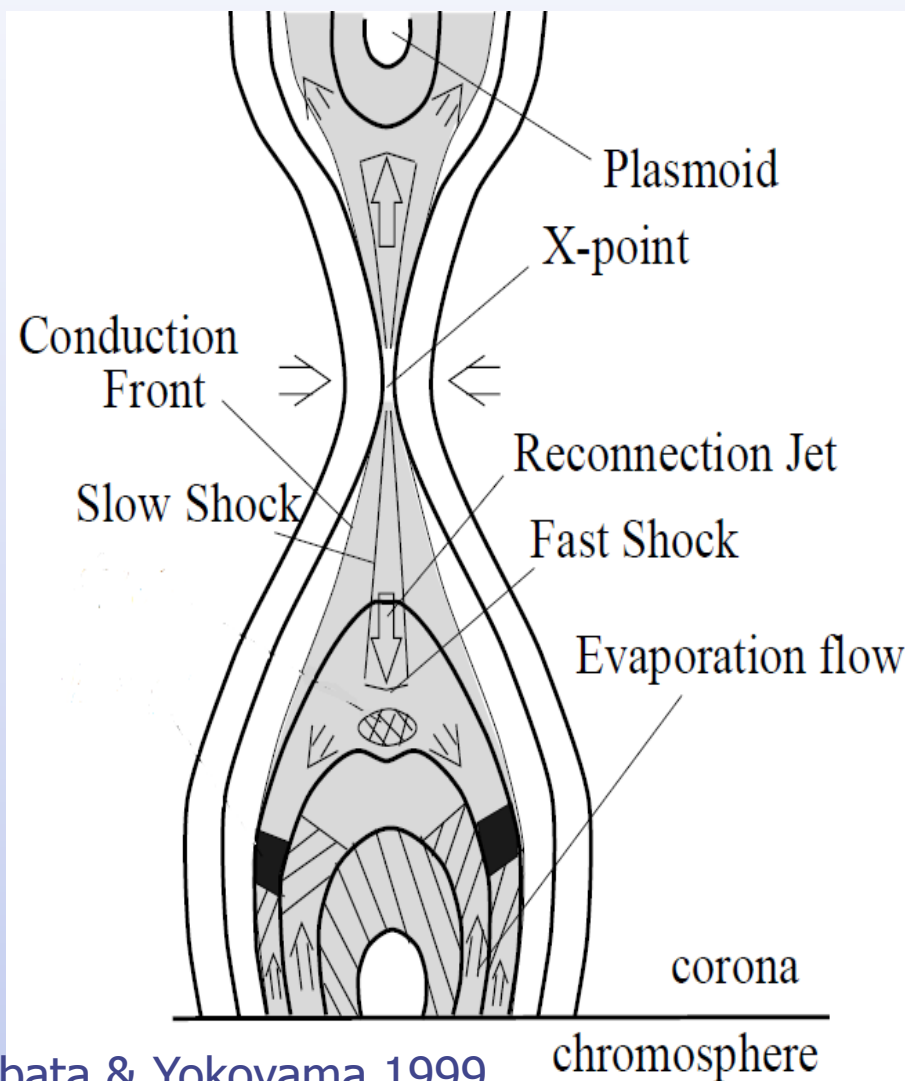


# Observations of coronal rain associated with solar flare by Hinode/SOT

Kyoko Watanabe  
ISAS/JAXA, Japan

# Basic Processes of Solar Flare

- ① Primary plasma heating & particle acceleration
  - Take place in the corona (reconnection points etc.)

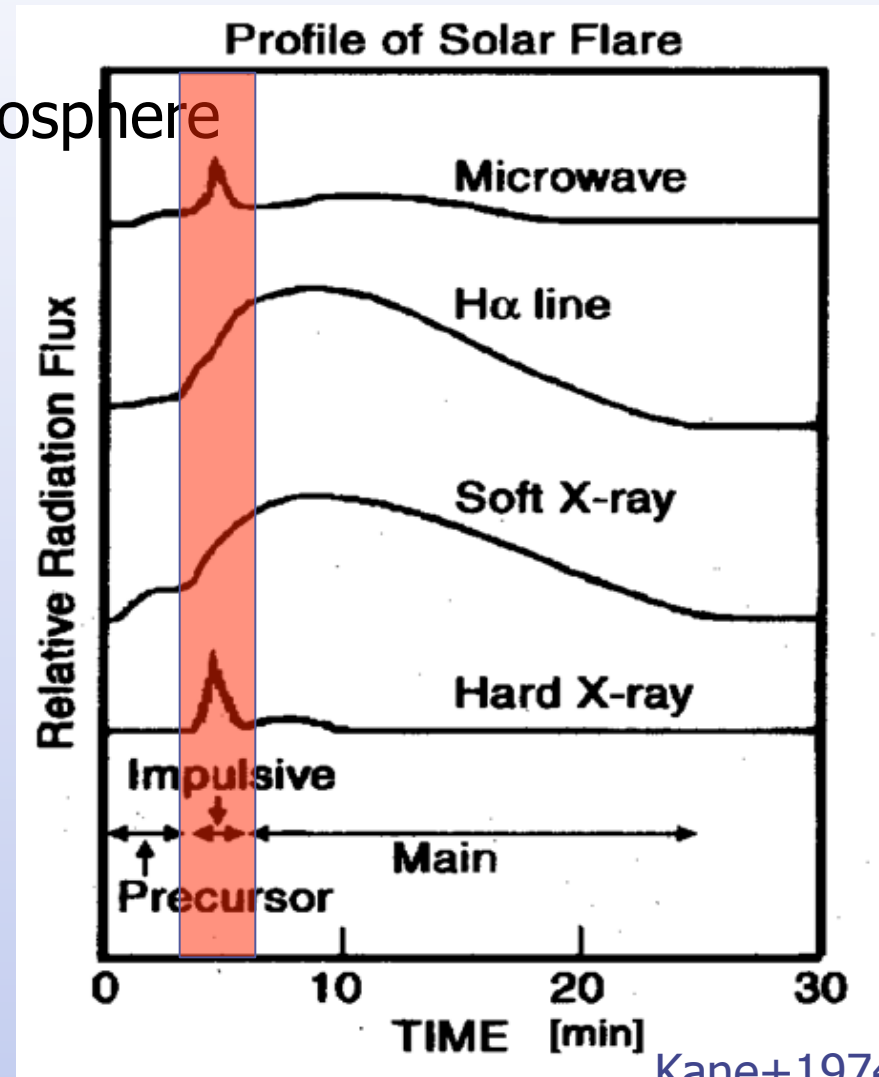
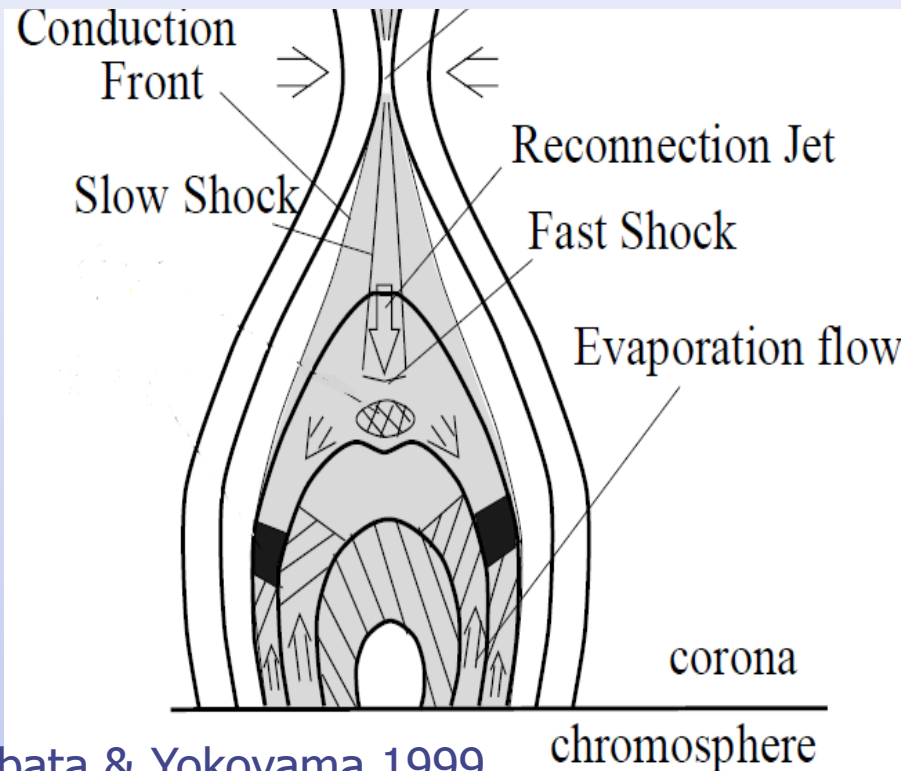


# Basic Processes of Solar Flare

- ① Primary plasma heating & particle acceleration
  - Take place in the corona (reconnection points etc.)

- ② Secondary heating process

- ① propagate to the chromosphere
- Observed in gamma-rays, HXRs, UV & WL

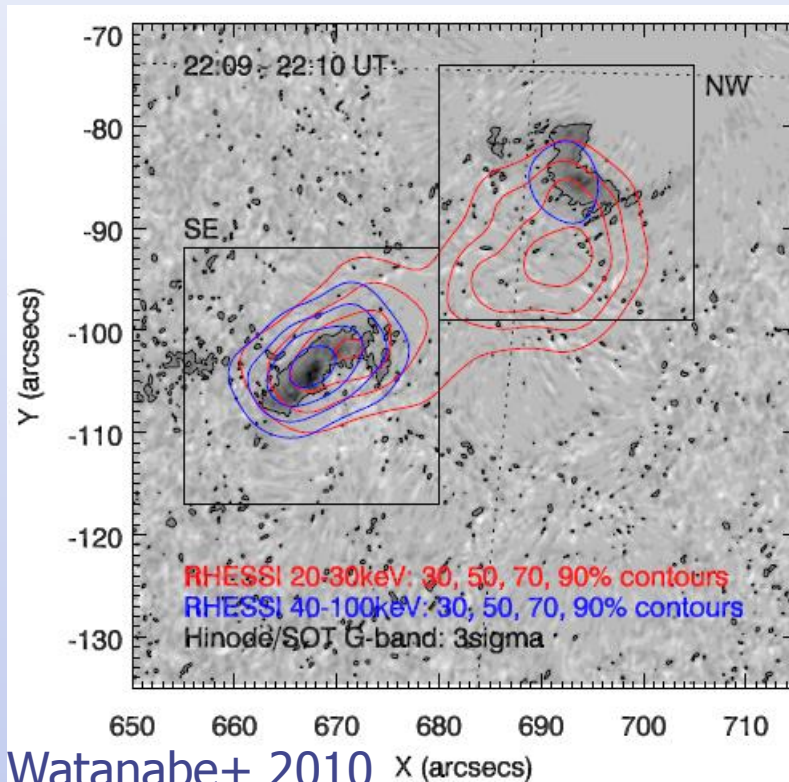


# Basic Processes of Solar Flare

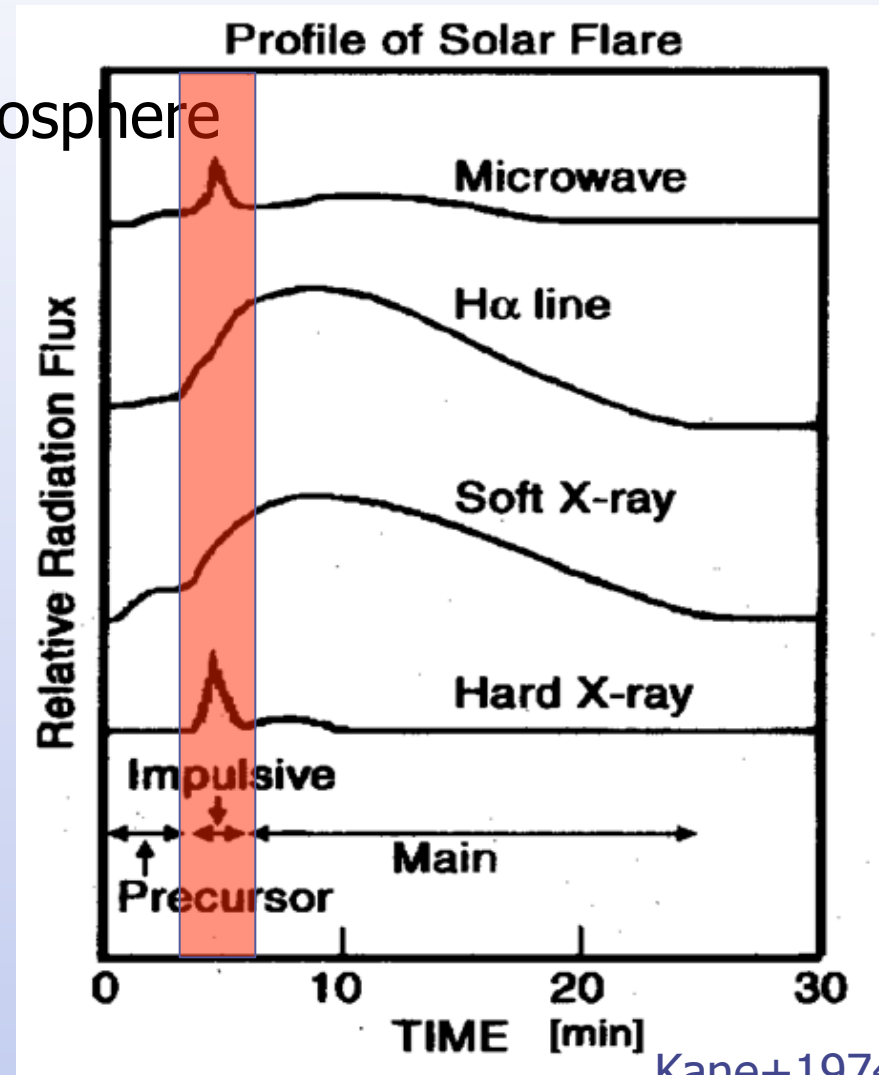
- ① Primary plasma heating & particle acceleration
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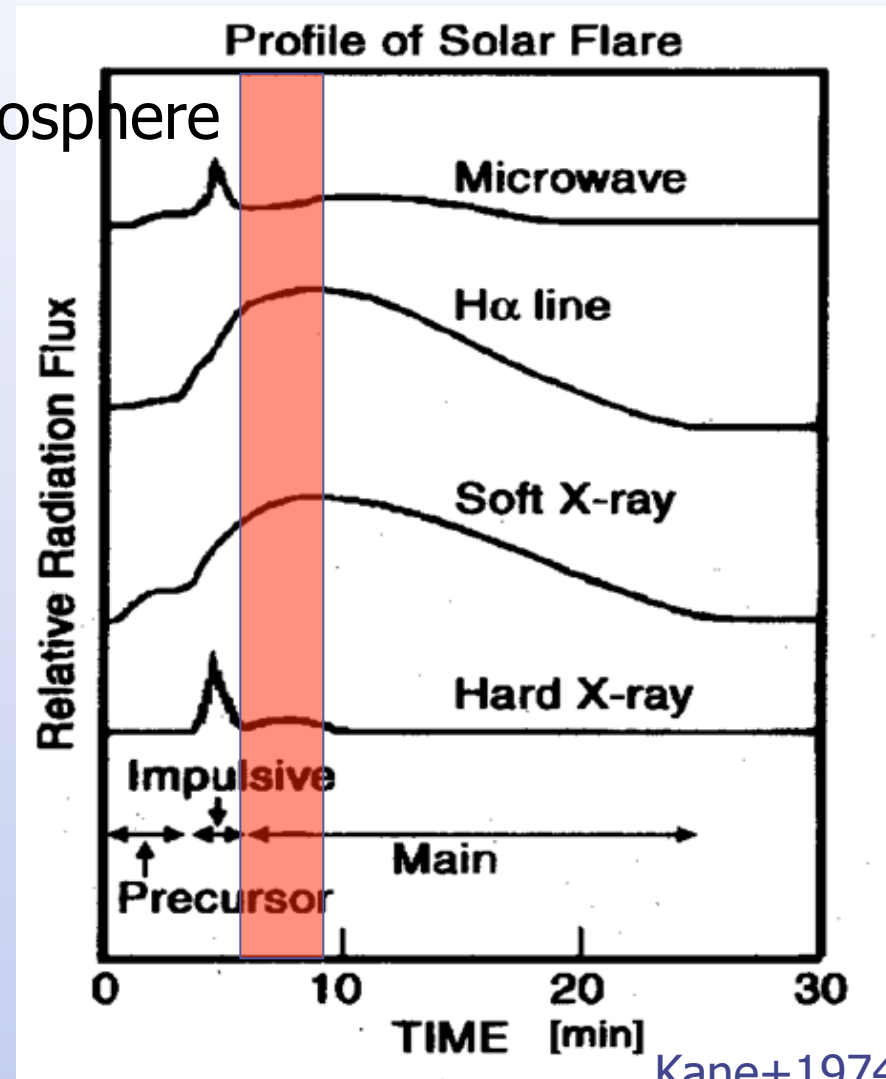
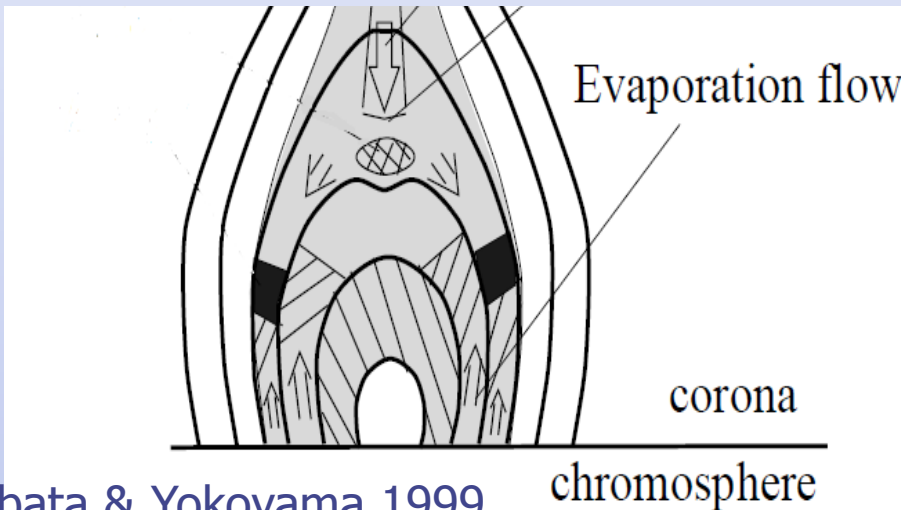
Watanabe+ 2010



Kane+1974

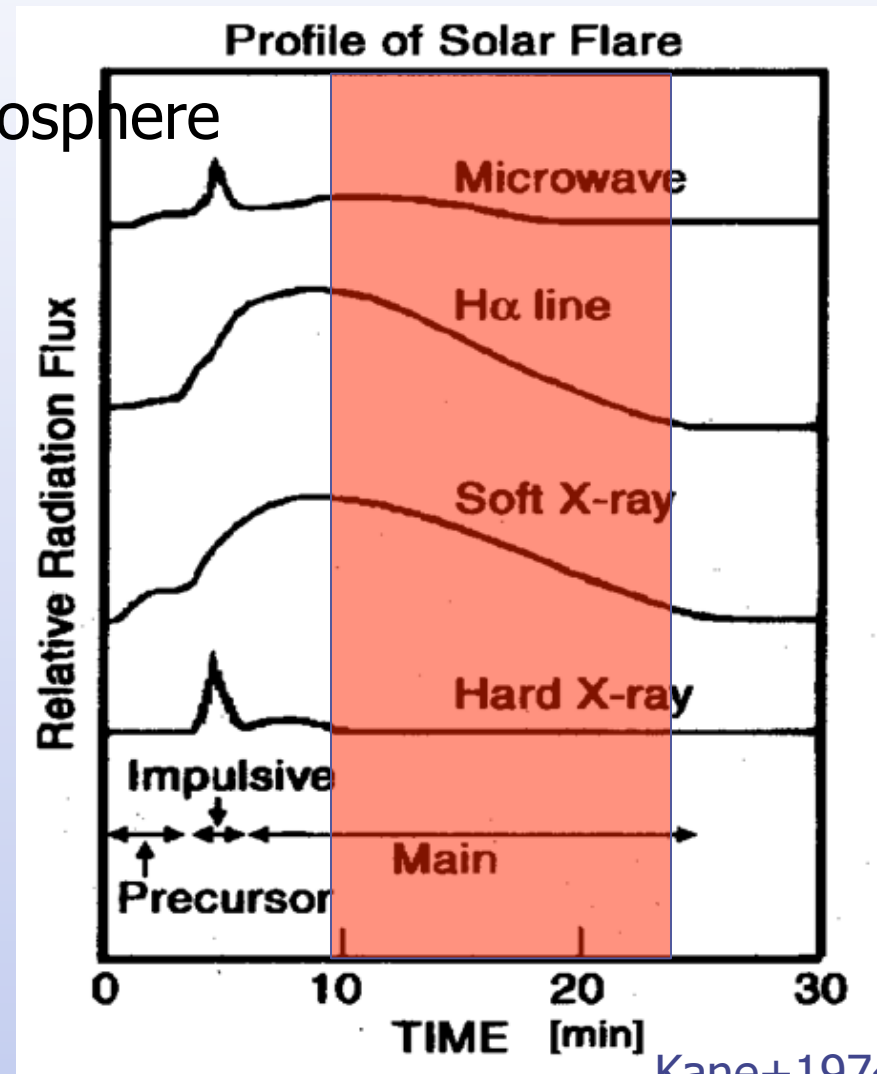
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- ③ Chromospheric evaporation
  - Formation of prominent flare loops in soft X-rays



# Basic Processes of Solar Flare

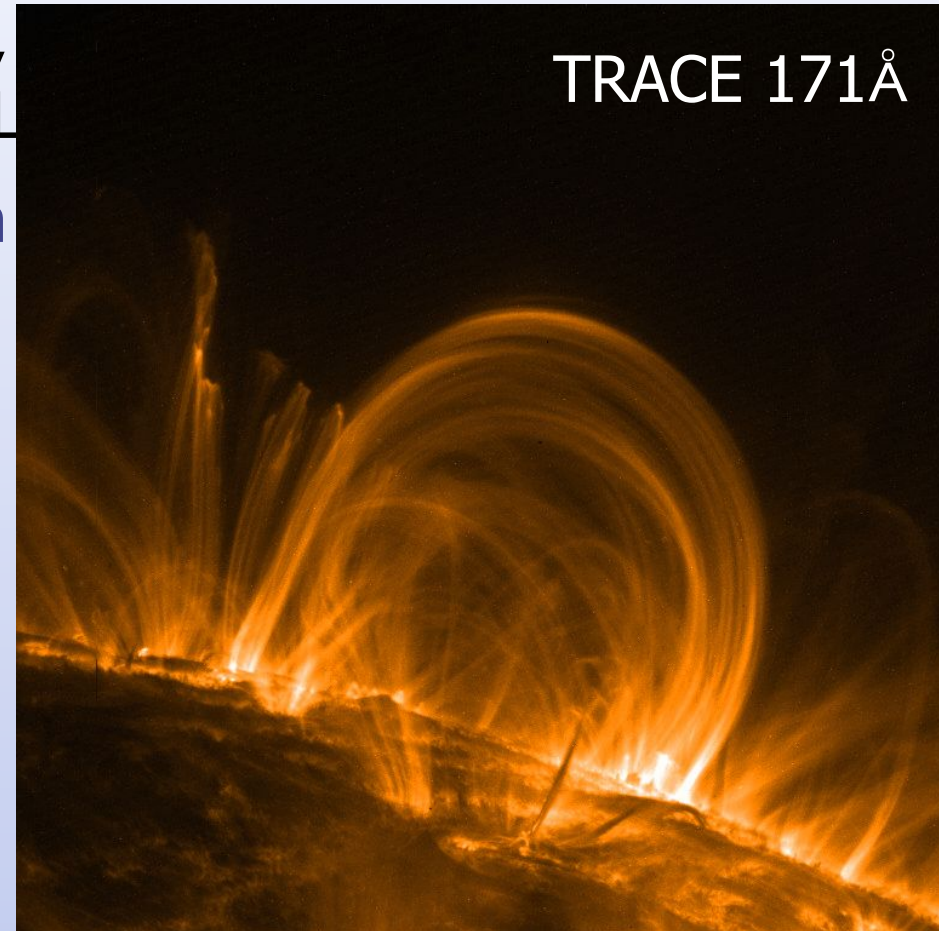
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  - Take place in the corona (reconnection points etc.)
- ② Secondary heating process
  - ① propagate to the chromosphere
  - Observed in gamma-rays, HXRs, UV & WL
- ③ Chromospheric evaporation
  - Formation of prominent flare loops in soft X-rays
- ④ Plasma cooling > heating
  - 10-30MK  $\rightarrow$  1-3MK
  - Post-flare loops become detectable in EUVs





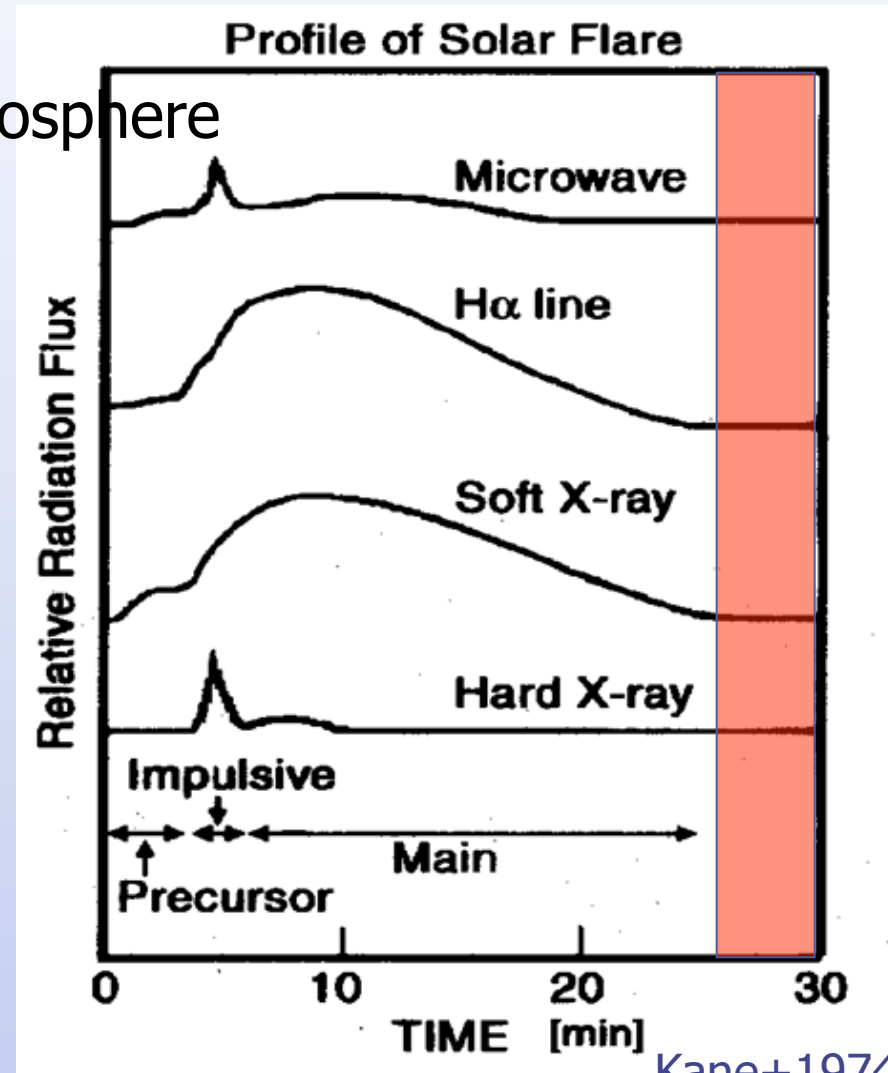
# Basic Processes of Solar Flare

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# Basic Processes of Solar Flare

- ① Primary plasma heating & particle acceleration
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  - ① propagate to the chromosphere
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  - Formation of prominent flare loops in soft X-rays
- ④ Plasma cooling > heating
  - 10-30MK  $\rightarrow$  1-3MK
  - Post-flare loops become detectable in EUVs
- ⑤ Loop decay & precipitation
  - Visible in UV & H $\alpha$





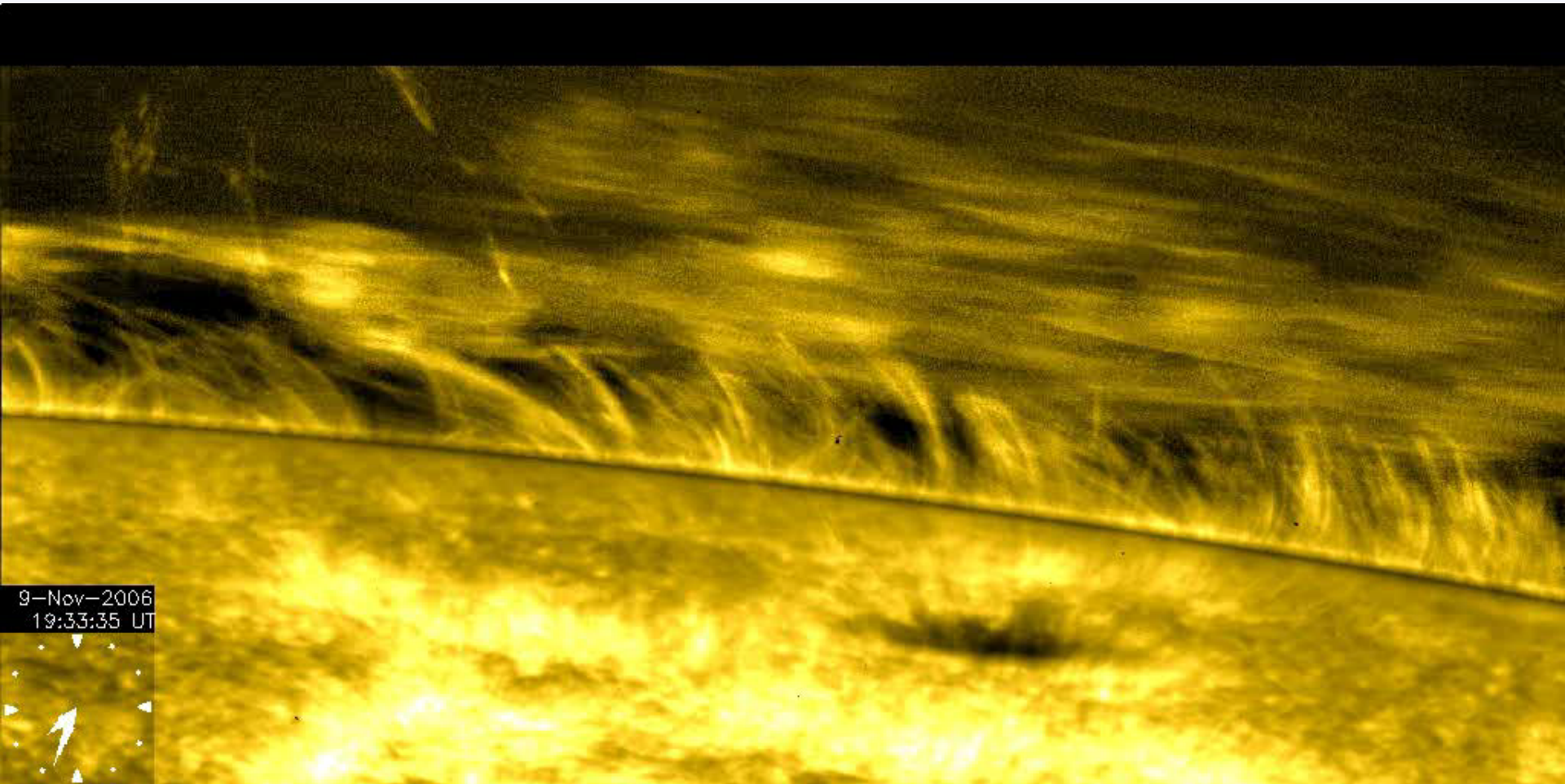
# Basic Processes of Solar Flare

- ① Primary plasma heating & particle acceleration
  - Take place in the corona (reconnection points etc.)
- ② Secondary heating process
  - ① propagate to the chromosphere ← energy injection to the chromosphere
  - Observed in gamma-rays, HXR's, UV & WL
- ③ Chromospheric evaporation
  - Formation of prominent flare loops in soft X-rays ← providing materials to the corona
- ④ Plasma cooling > heating
  - 10-30MK → 1-3MK ← coronal rain
  - Post-flare loops become detectable in EUVs
- ⑤ Loop decay & precipitation
  - Visible in UV & H $\alpha$

“Physics of the Solar Corona”  
2005, Markus Aschwanden  
Chapter 16:  
Flare Plasma Dynamics

# Coronal rain observations with Hinode/SOT

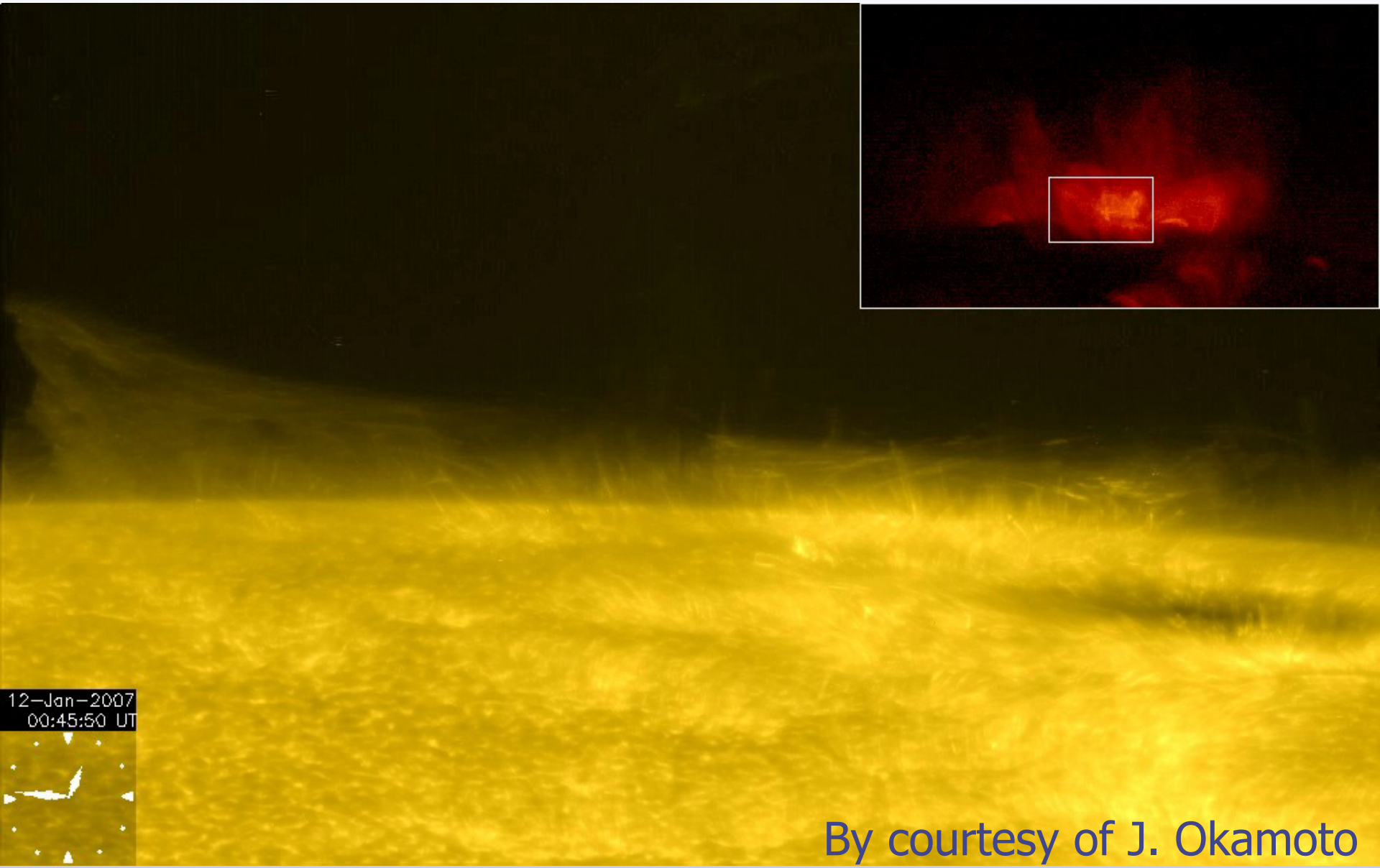
## Prominence & Coronal rain



By courtesy of J. Okamoto

# Coronal rain observations with Hinode/SOT

## C-class flare & Coronal rain

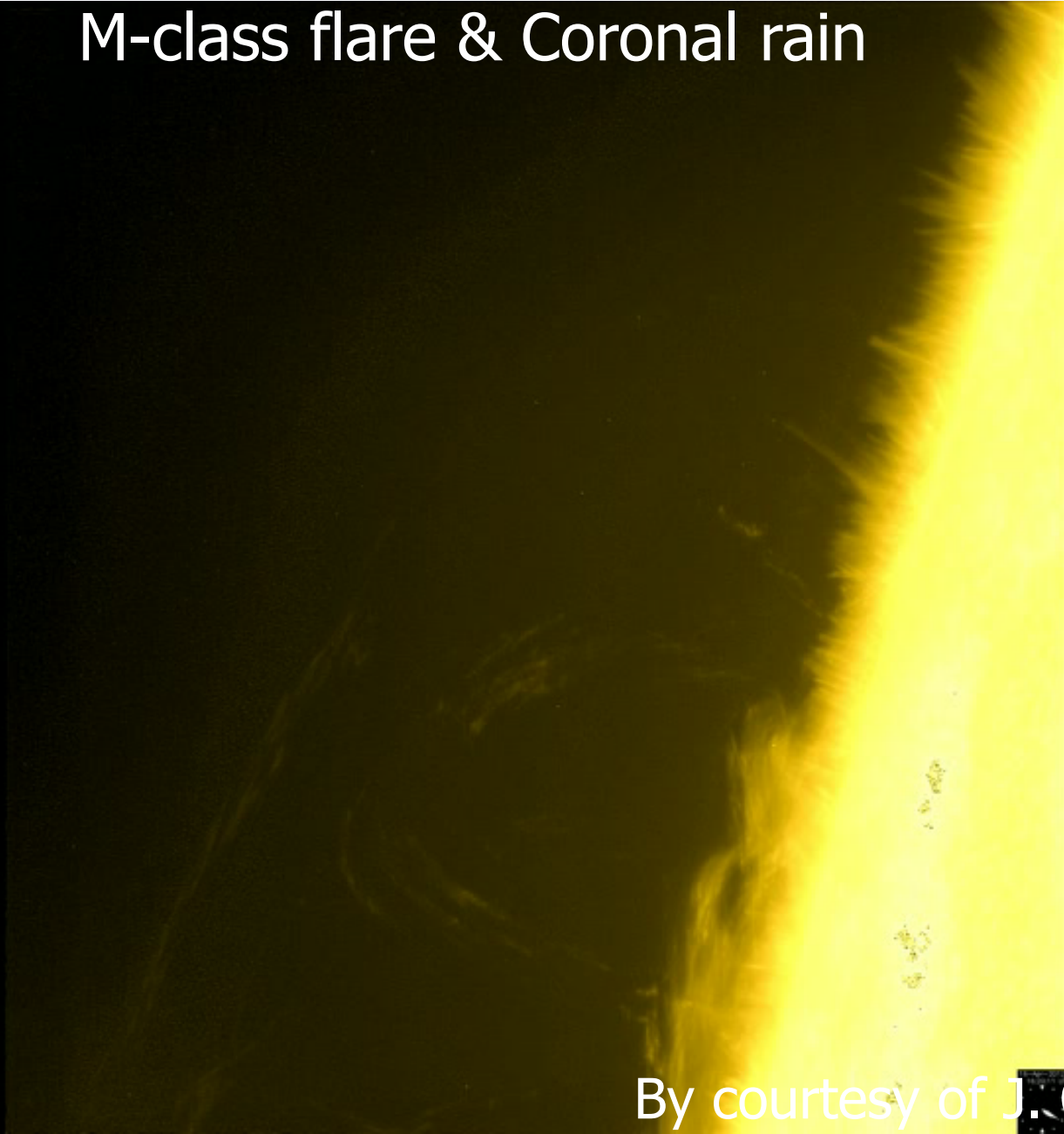


By courtesy of J. Okamoto



# Coronal rain observations with Hinode/SOT

## M-class flare & Coronal rain



By courtesy of J. Okamoto

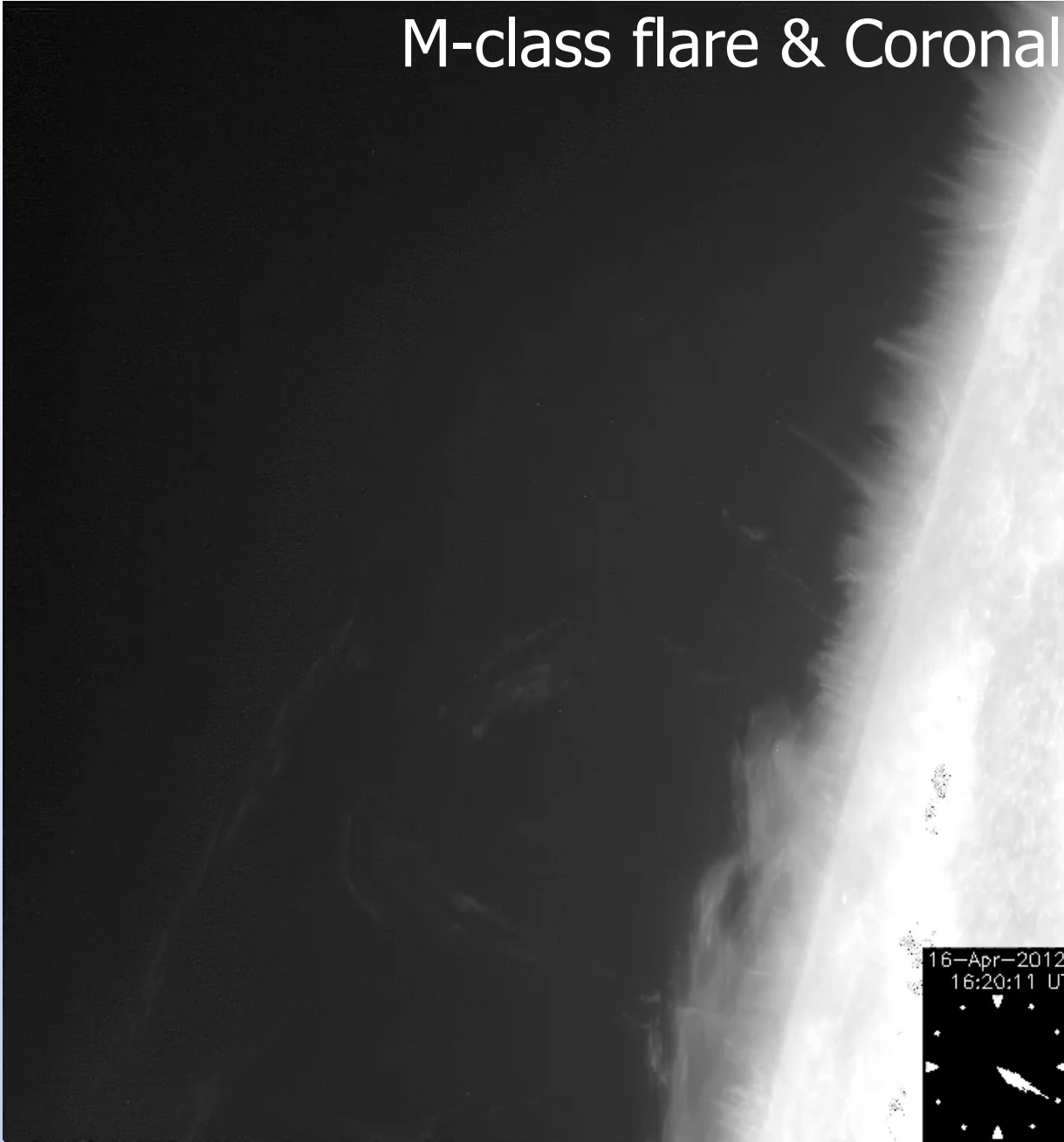
# Coronal rain observations with Hinode/SOT

## M-class flare & Coronal rain

2012 Apr 16  
M1.7-class flare

### Falling Materials

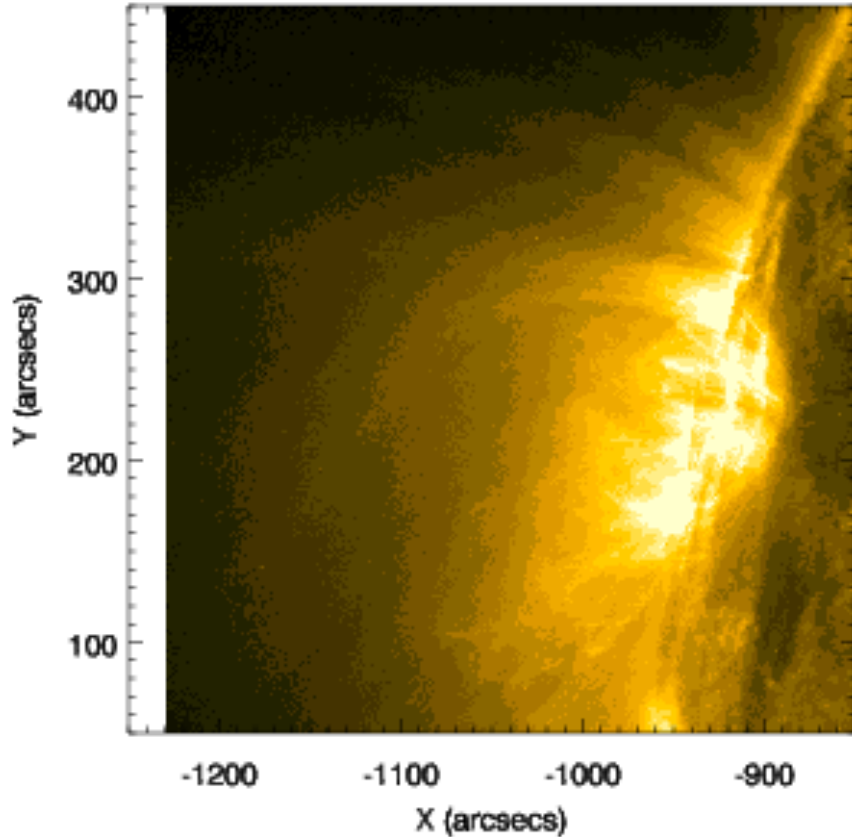
1. pre-flare coronal rain
2. quick fall down of flare plasma
3. prominence falling
4. post-flare loops



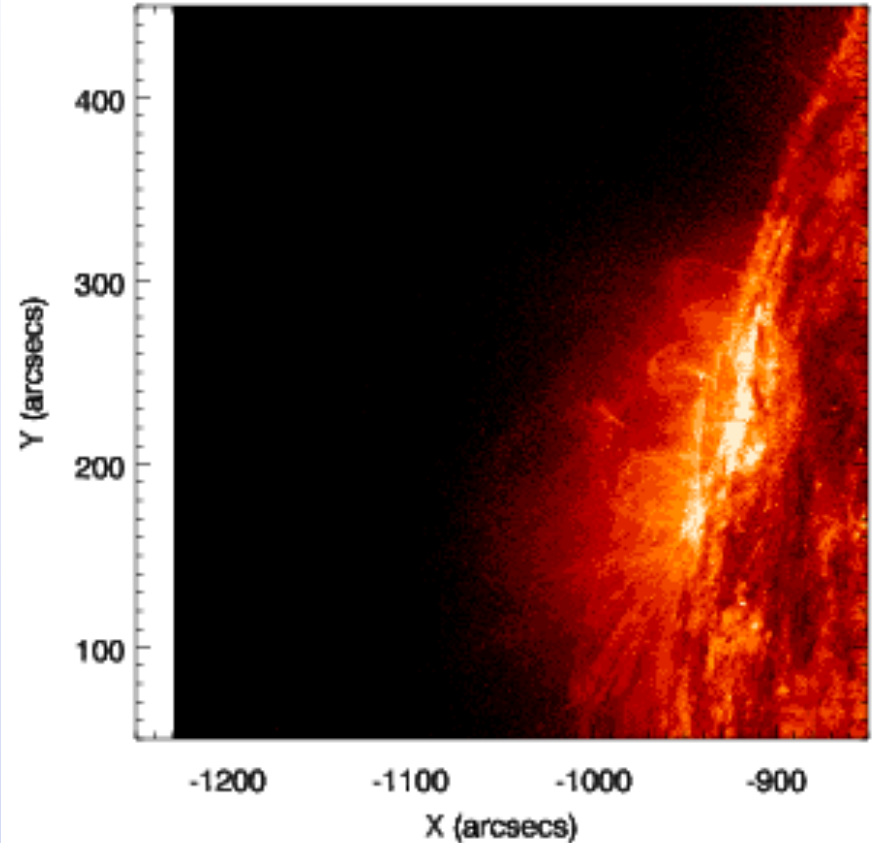
16-Apr-2012  
16:20:11 UT

# Coronal rain observations with SDO/AIA

SDO AIA\_3 171 16-Apr-2012 16:00:00.340 UT



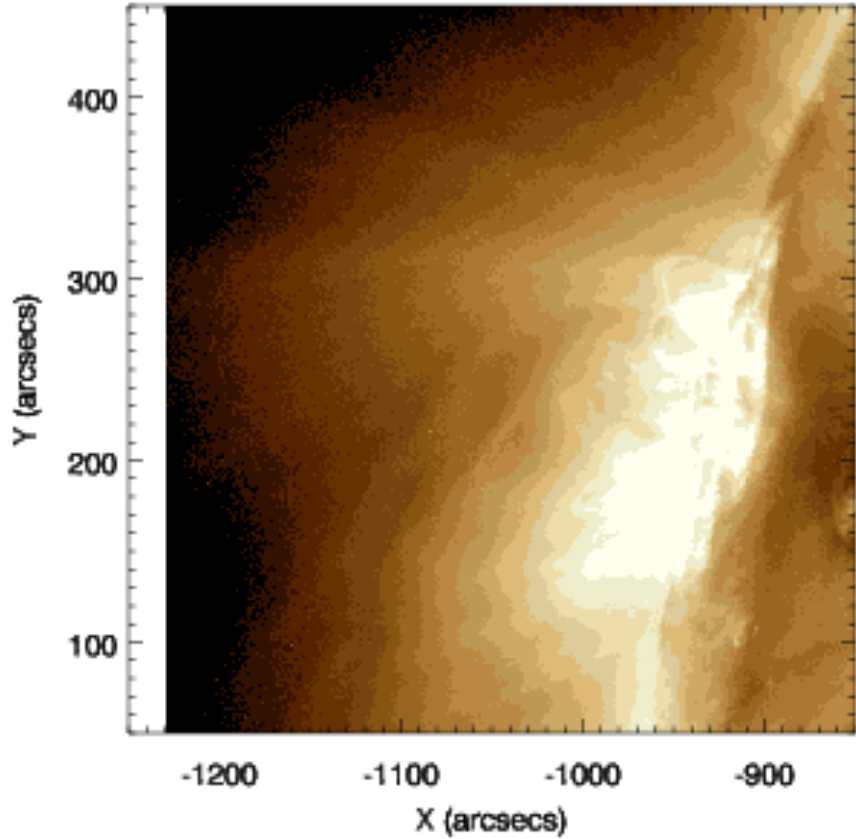
SDO AIA\_4 304 16-Apr-2012 16:00:08.120 UT



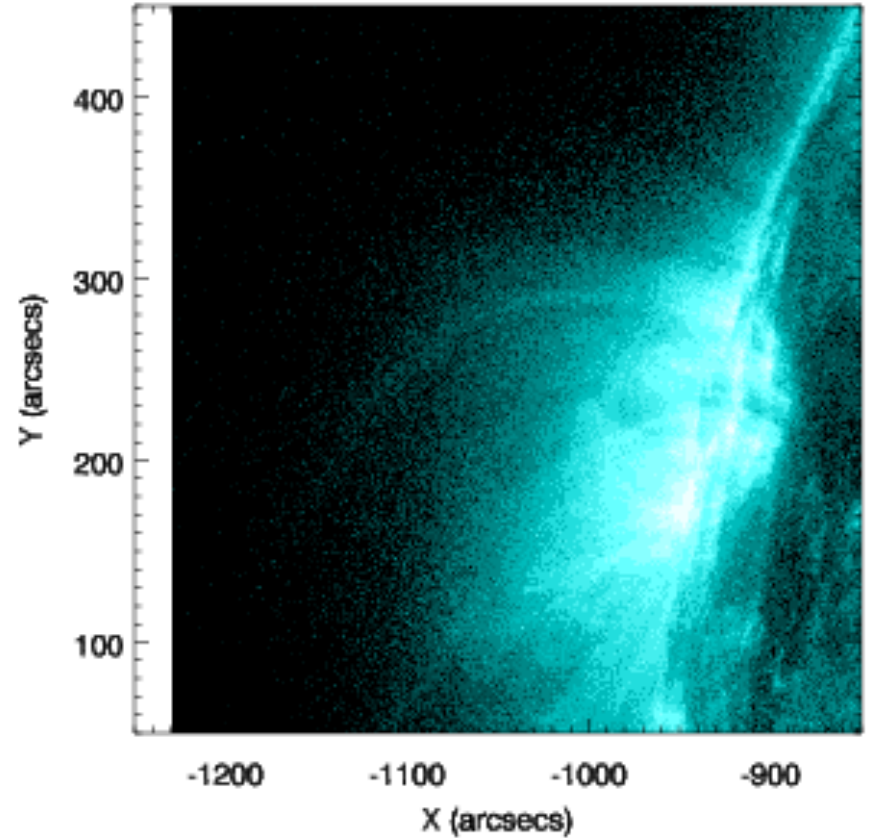


# Coronal rain observations with SDO/AIA

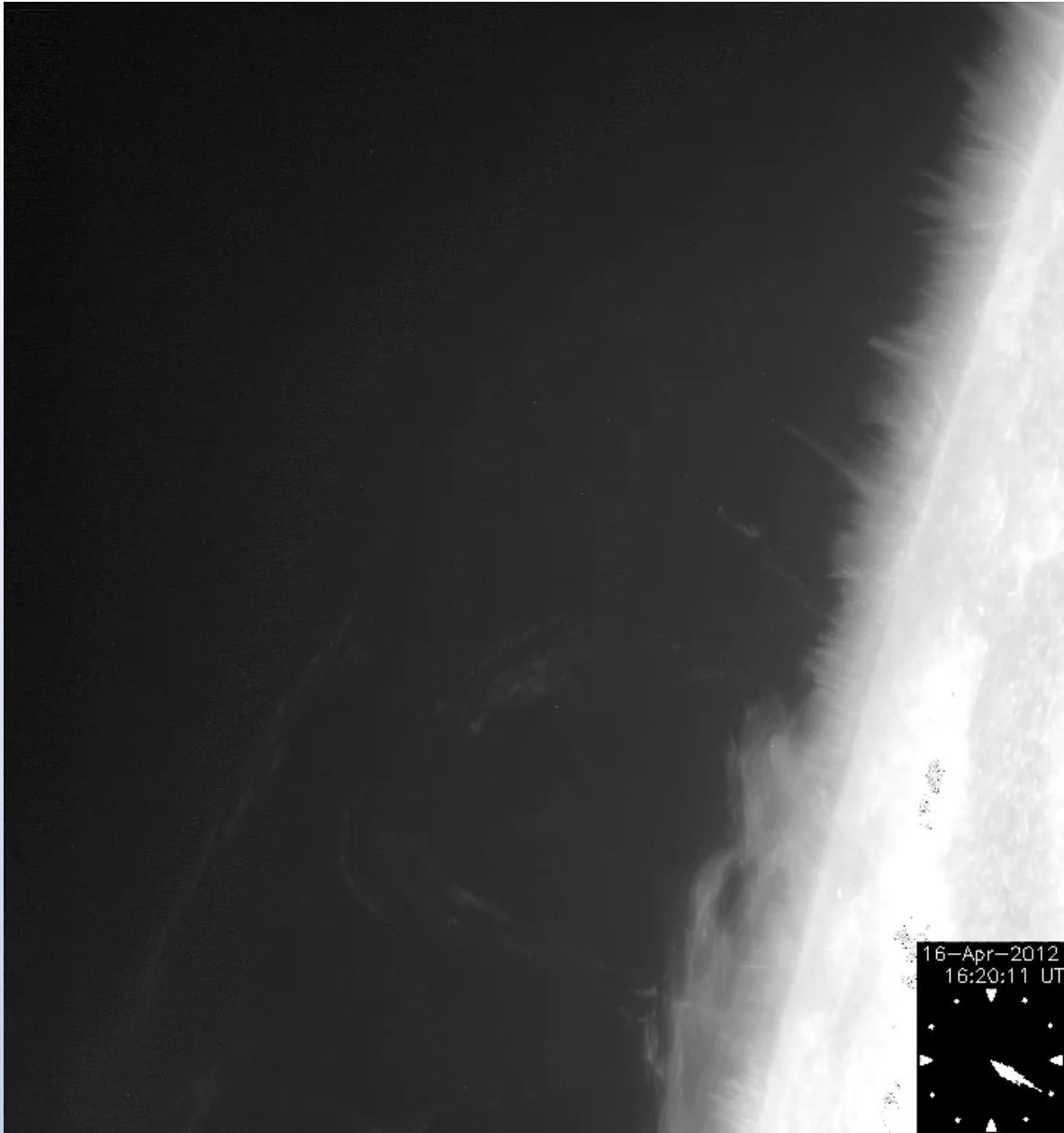
SDO AIA\_2 193 16-Apr-2012 16:00:43.840 UT



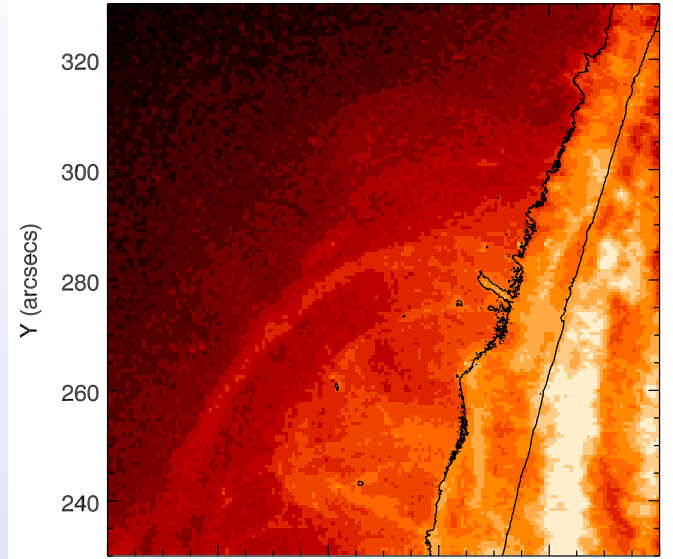
SDO AIA\_1 131 16-Apr-2012 16:00:33.620 UT



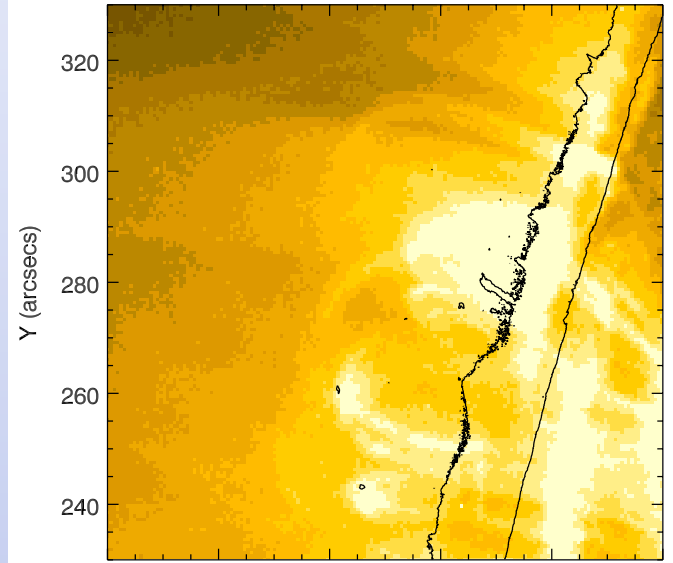
# Pre-Flare Coronal Rain



SDO AIA\_4 304 16-Apr-2012 16:30:32.120 UT

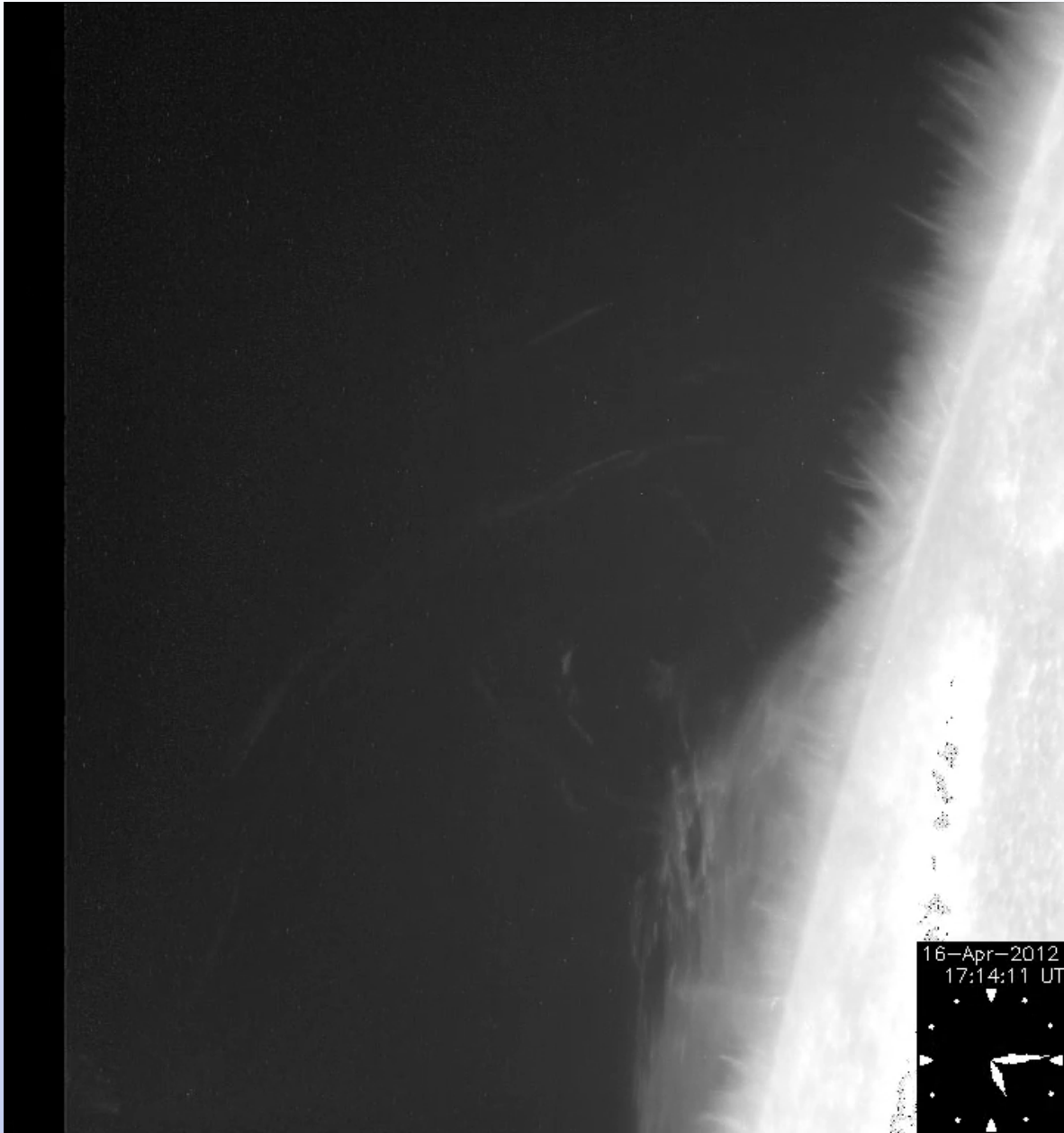


SDO AIA\_3 171 16-Apr-2012 16:30:12.340 UT

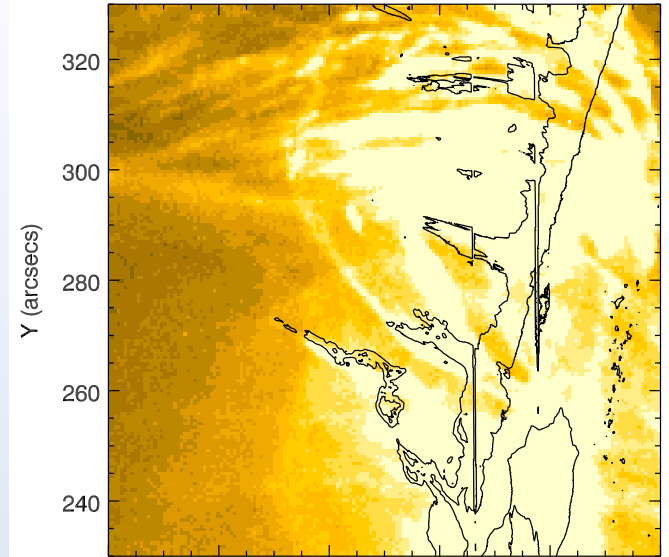


X (arcsecs)

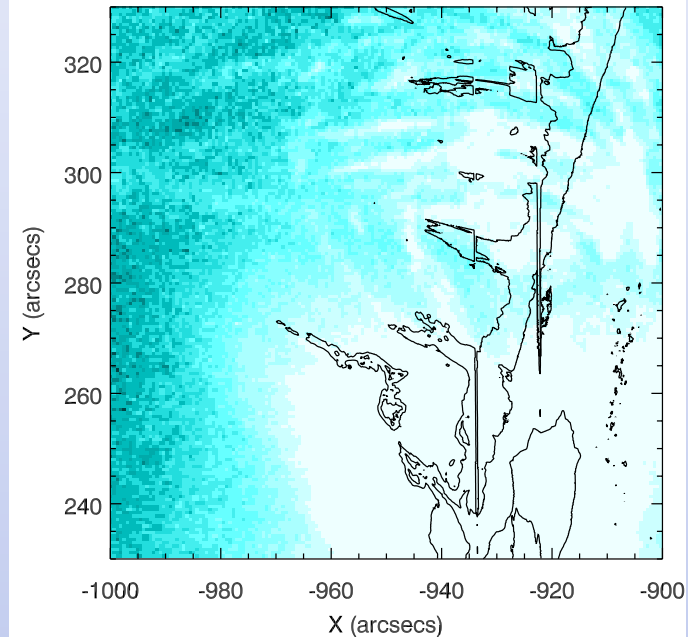
# Flaring - Flare Plasma Falling



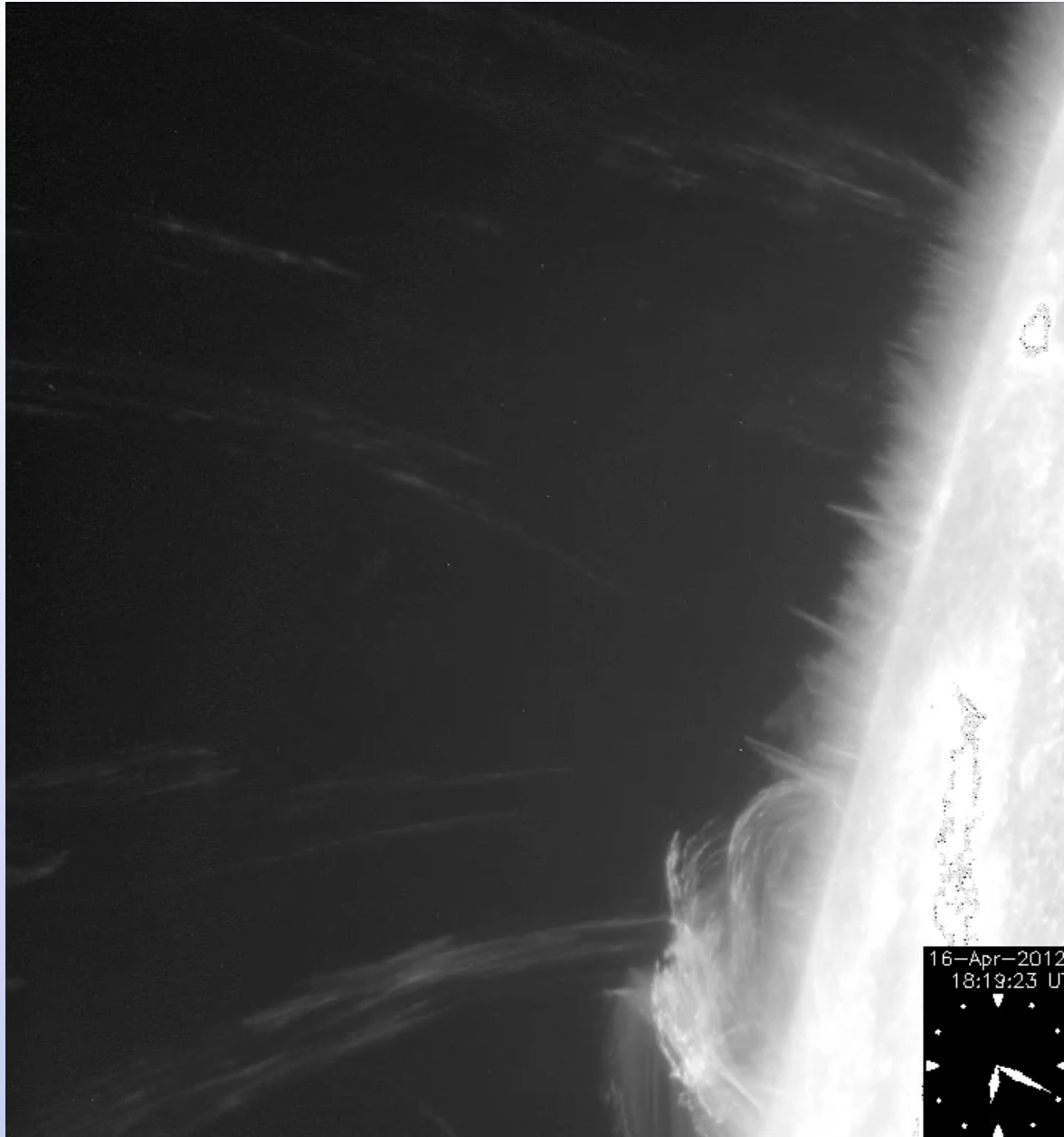
SDO AIA\_3 171 16-Apr-2012 17:46:14.300 UT



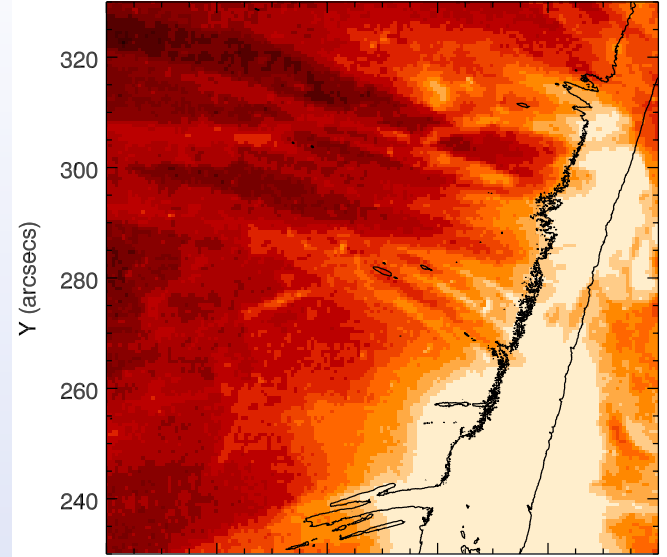
SDO AIA\_1 131 16-Apr-2012 17:46:12.350 UT



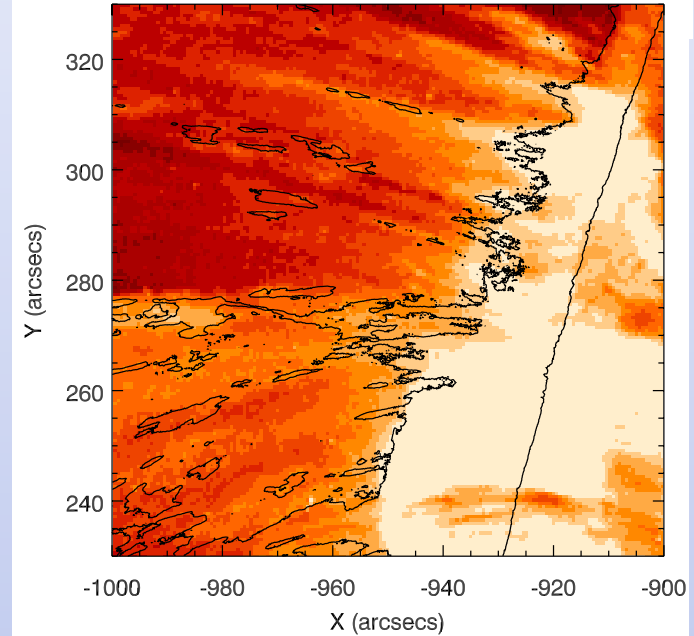
# Prominence Falling



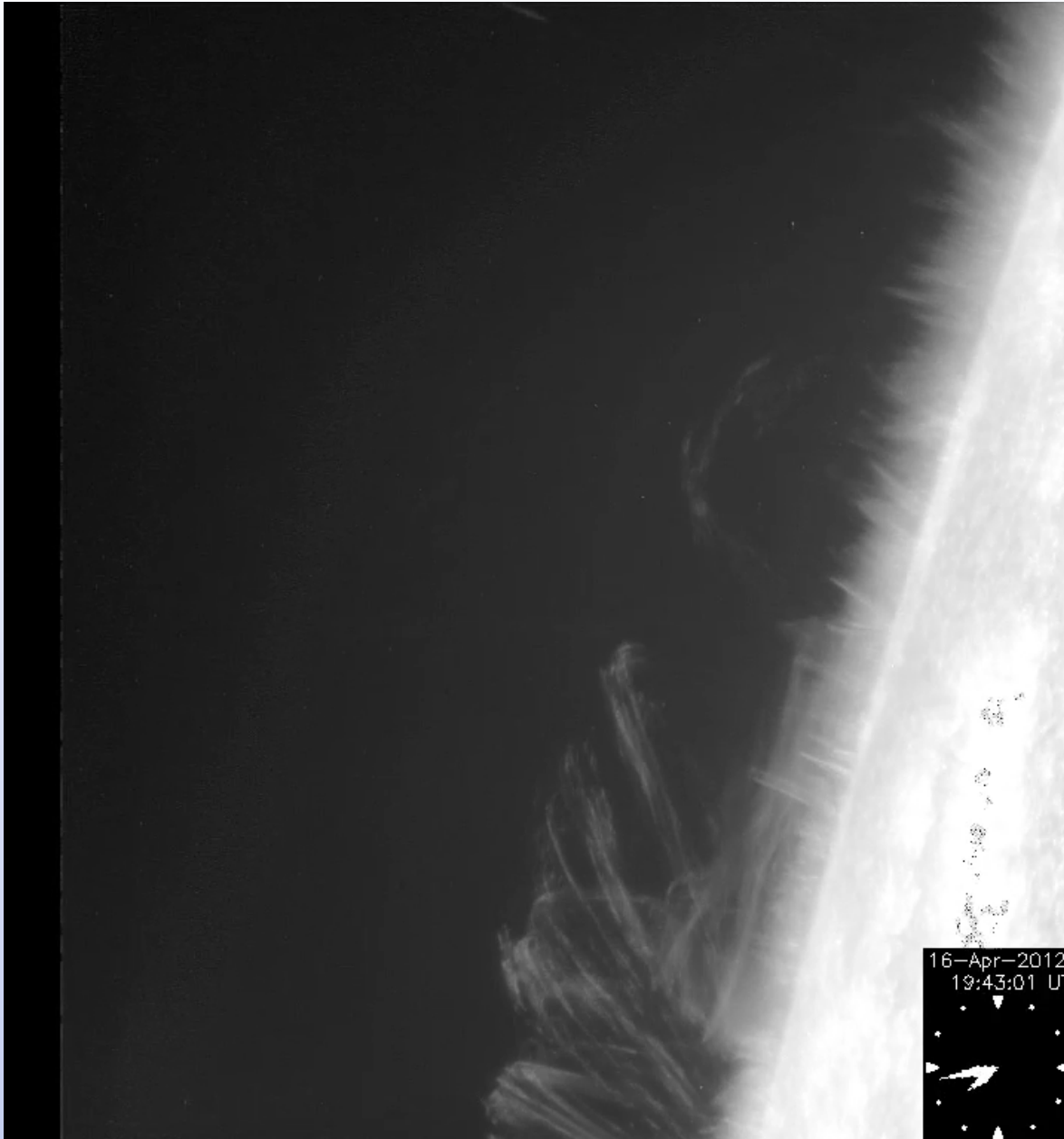
SDO AIA\_4 304 16-Apr-2012 18:20:08.120 UT



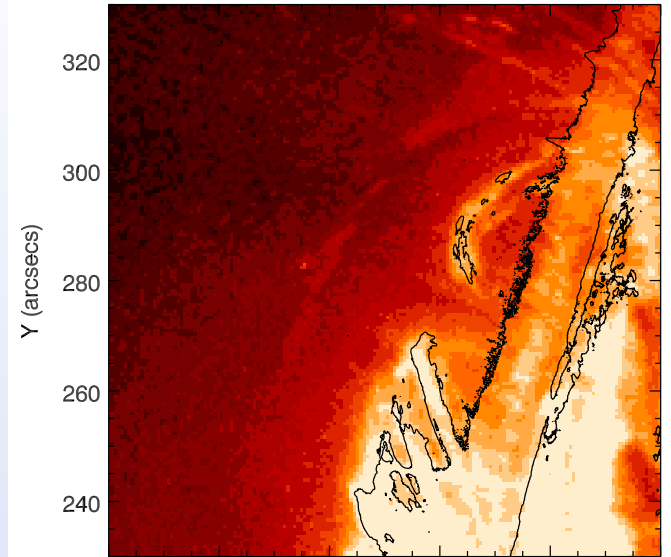
SDO AIA\_4 304 16-Apr-2012 18:58:08.130 UT



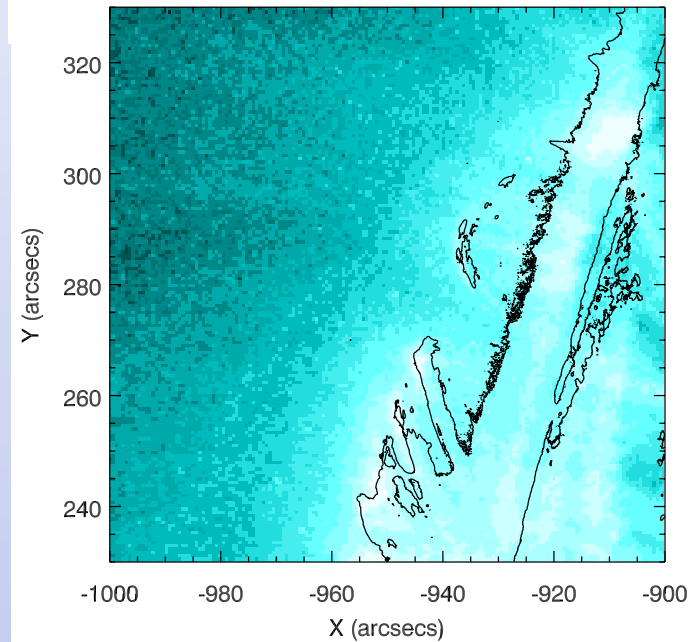
# Post-flare Loops



SDO AIA\_4 304 16-Apr-2012 19:46:08.120 UT



SDO AIA\_1 131 16-Apr-2012 19:40:57.620 UT



# Coronal rain candidates associated with solar flare observed by Hinode/SOT

| GOES start       | AR location | GOES | SOT |
|------------------|-------------|------|-----|
| 2007/06/02 10:28 | S04E74      | M1.0 | 63  |
| 2011/11/02 21:52 | N20E77      | M4.3 | 152 |
| 2011/11/03 10:58 | N20E70      | M2.5 | 87  |
| 2012/01/27 17:37 | N33W85      | X1.7 | 381 |
| 2012/04/16 17:24 | N14E88      | M1.7 | 92  |
| 2013/03/21 21:42 | N09W88      | M1.6 | 59  |
| 2013/06/07 22:11 | S32W89      | M5.9 | 196 |
| 2013/10/26 19:24 | S09E81      | M3.1 | 22  |
| 2014/03/13 19:03 | N15W87      | M1.2 | 203 |
| 2014/05/06 22:01 | S10W87      | M1.0 | 104 |
| 2014/10/29 09:54 | S18W77      | M1.2 | 160 |
| 2014/10/29 14:19 | S16W81      | M1.4 | 245 |
| 2014/10/29 16:06 | S14W82      | M1.0 | 140 |
| 2014/10/29 18:47 | S13W77      | M1.3 | 30  |
| 2014/10/30 00:34 | S14W81      | M1.3 | 35  |
| 2014/10/30 01:19 | S14W88      | M3.5 | 190 |
| 2014/10/30 04:17 | S16W89      | M1.2 | 100 |

## Selection condition

- Limb flare ( $>70\text{deg}$ )
- $\geq$  M-class flare
- SOT data existence  $>10$  images

## Green events

- flare mode obs. images available every 20 sec for each band (Ca, RGB)
- NOAA12192: 7 events