

Few words on g-mode asymptotic

T. Corbard, J. Provost, L. Bigot

1 Asymptotic model Provost & Berthomieu (1986)

Their Eq.(4a) writes:

$$P_{n,l} = \frac{P_0}{\sqrt{l(l+1)}} \left[n + \frac{l}{2} - \frac{1}{4} - \theta \right] + \frac{P_0^2}{P_{n,l}} \frac{l(l+1)V_1 + V_2}{l(l+1)}$$

- Use Cowling approximation
- θ is not a constant but is function of the frequency and l
- V_1 depends on the stratification near the center of the Sun
- V_2 and θ depends predominantly on the stratification and mixing processes at the base of the convection zone

→ Possible interplay between V_2 and θ

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$$LP \sim \left(|n| + \frac{1}{2}l + \alpha_g \right) P_0 - (A_g L^2 - B_g) \frac{P_0^2}{P} + \dots,$$

Scherrer & Gough L

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- V_2 and θ depends predominantly on the stratification and mixing processes at the base of the convection zone

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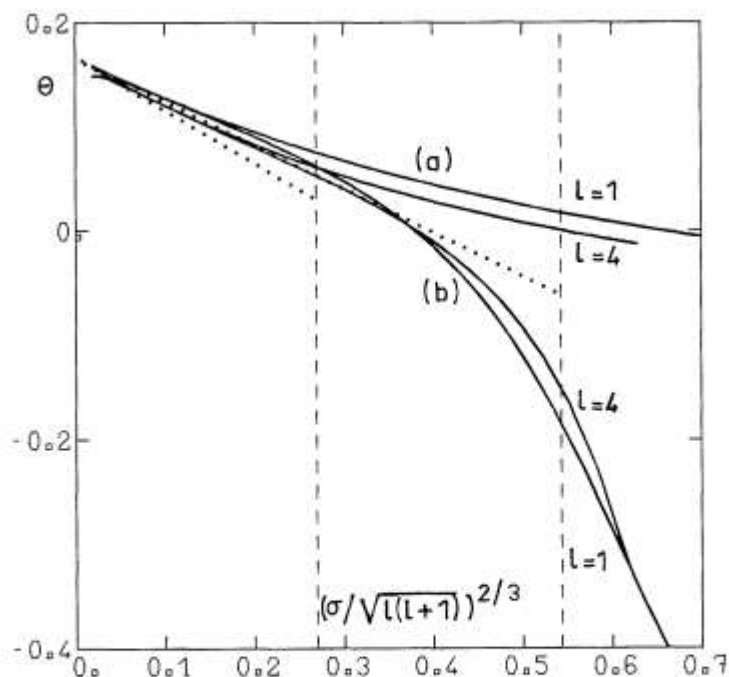


Fig. 2. Variation of θ relatively to $[\sigma^2 / l(l+1)]^{1/3}$ for the standard solar model and for $l = 1, 4$ at first and second order relatively to σ , respectively curves (a) and (b). The dotted lines represent the asymptotic behavior (formula 8). The dashed vertical lines delimit the range of nondimensional frequencies corresponding approximately to modes of radial order $10 \leq k \leq 35$, $\varphi_0 = 138$, $\varphi_1 = -6.3$

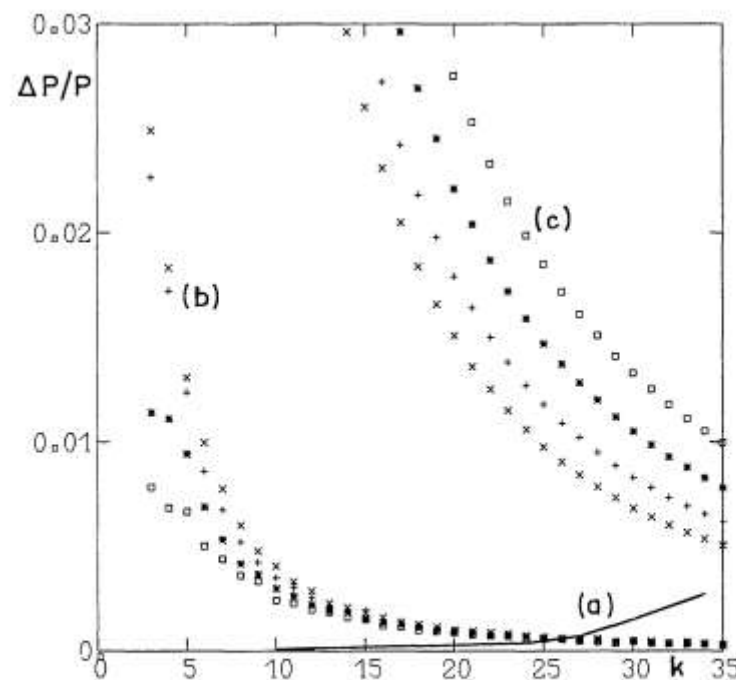


Fig. 4. Relative variation of the periods relatively to the radial order k , measuring: (a) the numerical errors; (b) the effect of the perturbation of gravitational potential; (c) the importance of second order terms in asymptotic formula(10) ($\times$ $l = 1$; $+$ $l = 2$; $*$ $l = 3$; $\square$ $l = 4$)

g-mode asymptotic : observations

2 Fossat et al. (2017)

Assuming that $n \gg 1$, $P_l = \frac{P_0}{\sqrt{l(l+1)}}$

$$P_{n,l} = P_{n_0,l} + (n - n_0) \left(P_l - \frac{\alpha}{nn_0} \right) + O\left(\frac{1}{n^3}\right) \quad \alpha = P_l(l(l+1)V_1 + V_2)$$

Fossat et al. 2017

$$\alpha_1 = 1350 \text{ s} \quad P_1 \simeq 1443 \text{ s} \Rightarrow 2V_1 + V_2 \simeq 0.94$$

Fossat & Schmider 2018

$$\begin{aligned} - P_{n,1} &= P_{-22,1} + (|n| - 22)(P_1 - 50/|n|) \text{ for } n = -95 \dots -23. \\ - P_{n,2} &= P_{-36,2} + (|n| - 36)(P_2 - 4/|n|) \text{ for } n = -142 \dots -37. \end{aligned}$$

$$\alpha_1 = 1100 \text{ s}$$

$$\alpha_2 = 144 \text{ s}$$

$\Rightarrow$

$$V_1 \simeq -0.147$$

$$V_2 \simeq 1.06$$

Asymptotic parameters from models

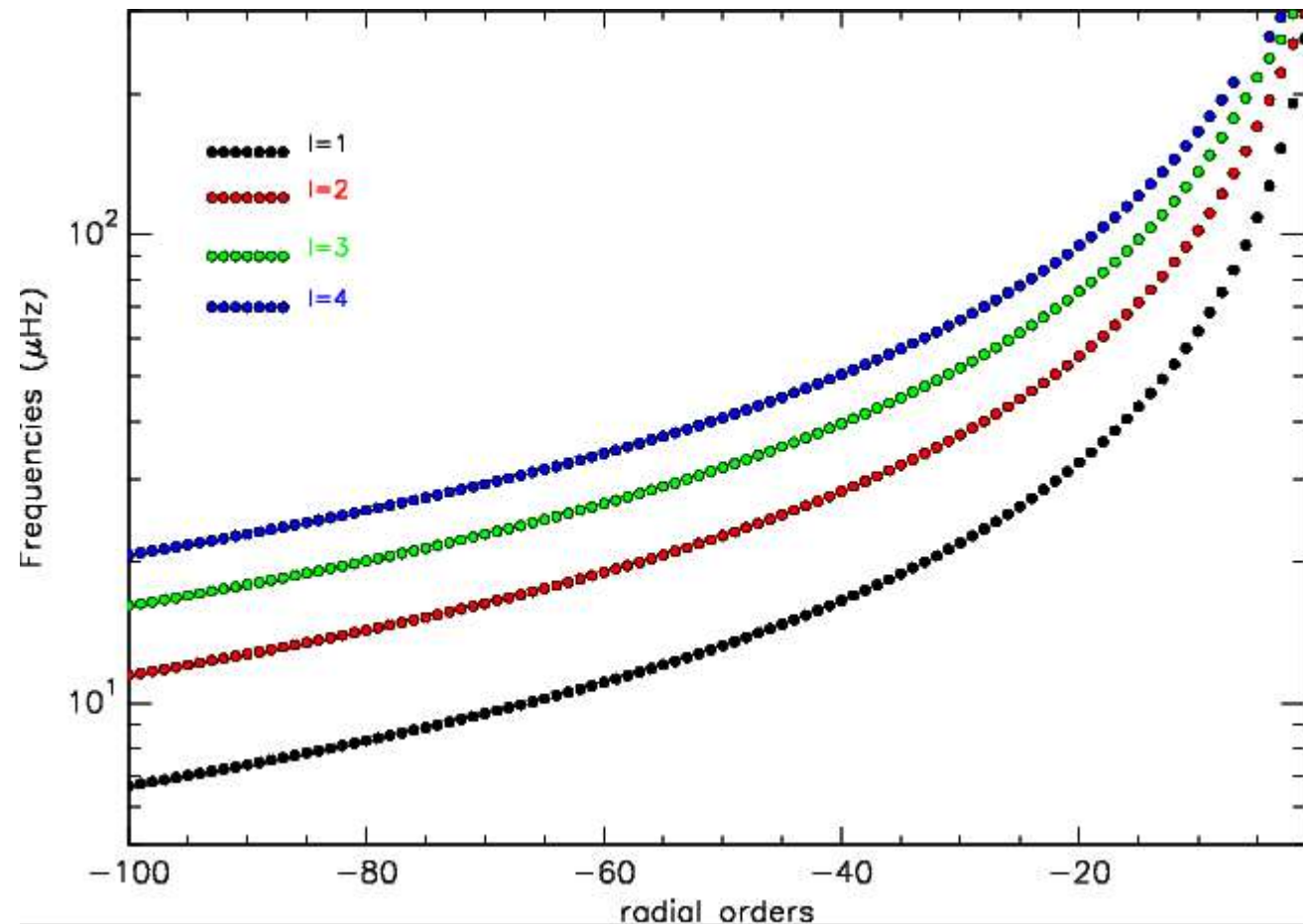
Two approaches :

(1) Fitting theoretical periods

=> model+oscillation code

(2) Directly from a model

Numerical frequencies : CESAM+GYRE

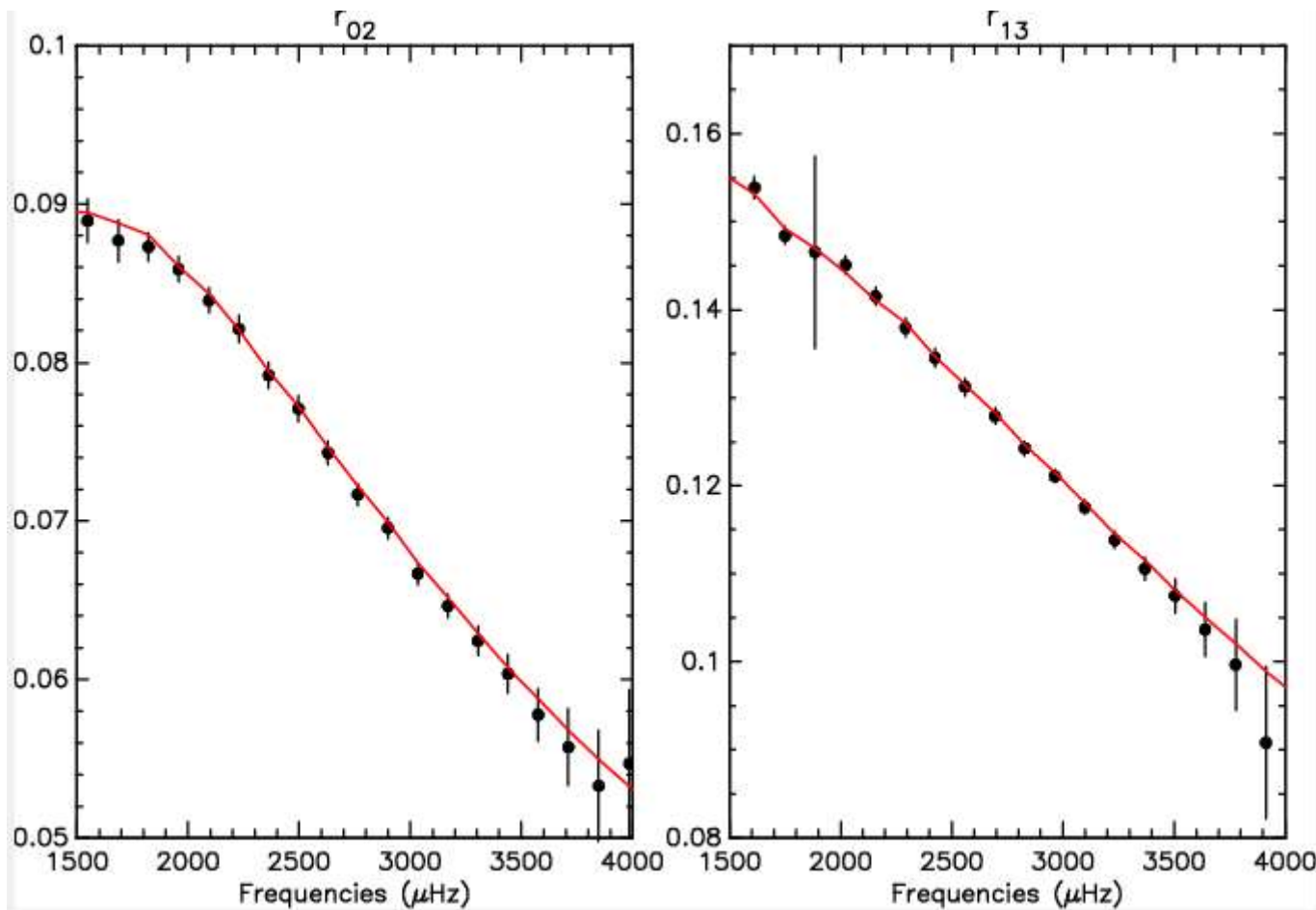


CESAM

$M = 1.988475000000\text{E}+33$ g
 $R = 6.958842022607\text{E}+10$ cm
 $L = 3.822689401918\text{E}+33$ erg/s
 $T_{\text{eff}} = 5774.2$ K
 $Y = 0.2423$
 $Z/X = 0.0244$
Age = 4.57 Gy
Base of cz = 0.712
 $P_0 = 35.34$ min

OPAL,
MLT Böhm-Vitesse ($\alpha=2.16$),
diffusion Michaud-Proffitt,
NACRE for nuclear reactions,
MARCS atmosphere,
no rotation,
no overshoot.

Numerical frequencies : CESAM+GYRE

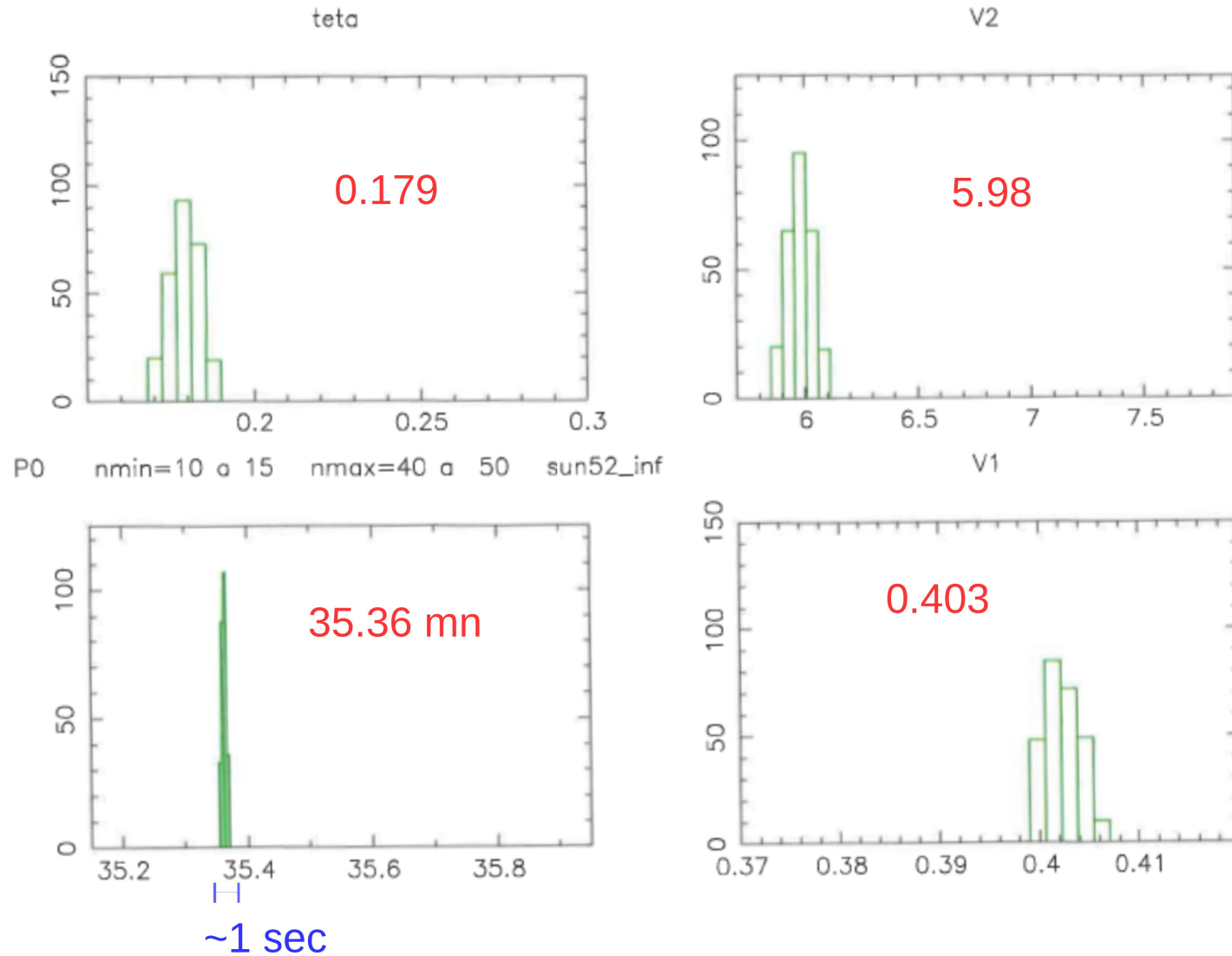


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Fitting the 2d order asymptotic to numerical frequencies



g-mode asymptotic : observations

2 Fossat et al. (2017)

Assuming that $n \gg 1$, $P_l = \frac{P_0}{\sqrt{l(l+1)}}$

$$P_{n,l} = P_{n_0,l} + (n - n_0) \left(P_l - \frac{\alpha}{nn_0} \right) + O\left(\frac{1}{n^3}\right) \quad \alpha = P_l(l(l+1)V_1 + V_2)$$

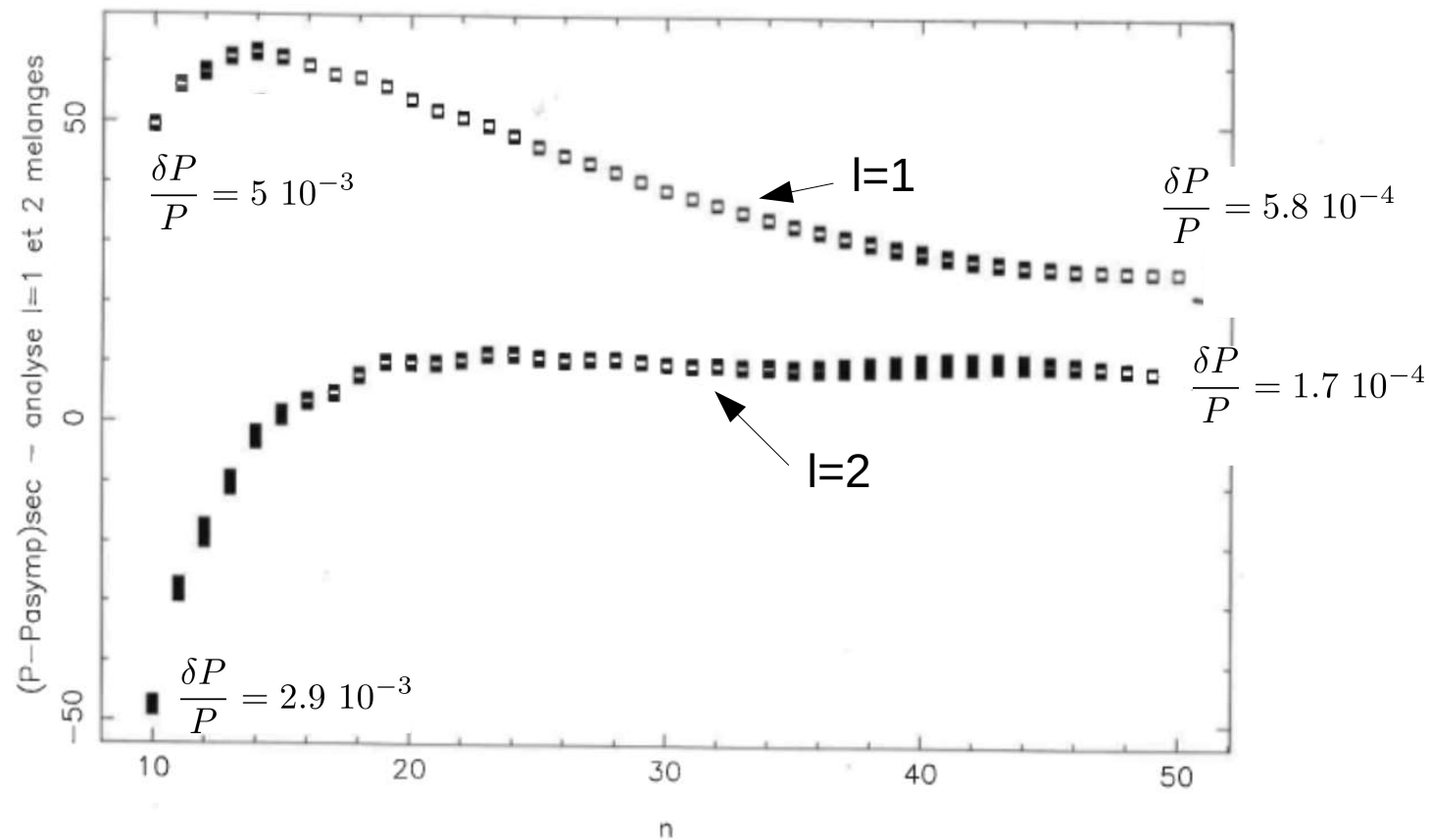
Fossat et al. 2017

$$\alpha_1 = 1350 \text{ s} \quad P_1 \simeq 1443 \text{ s} \Rightarrow 2V_1 + V_2 \simeq 0.94 \quad (\sim 6.8)$$

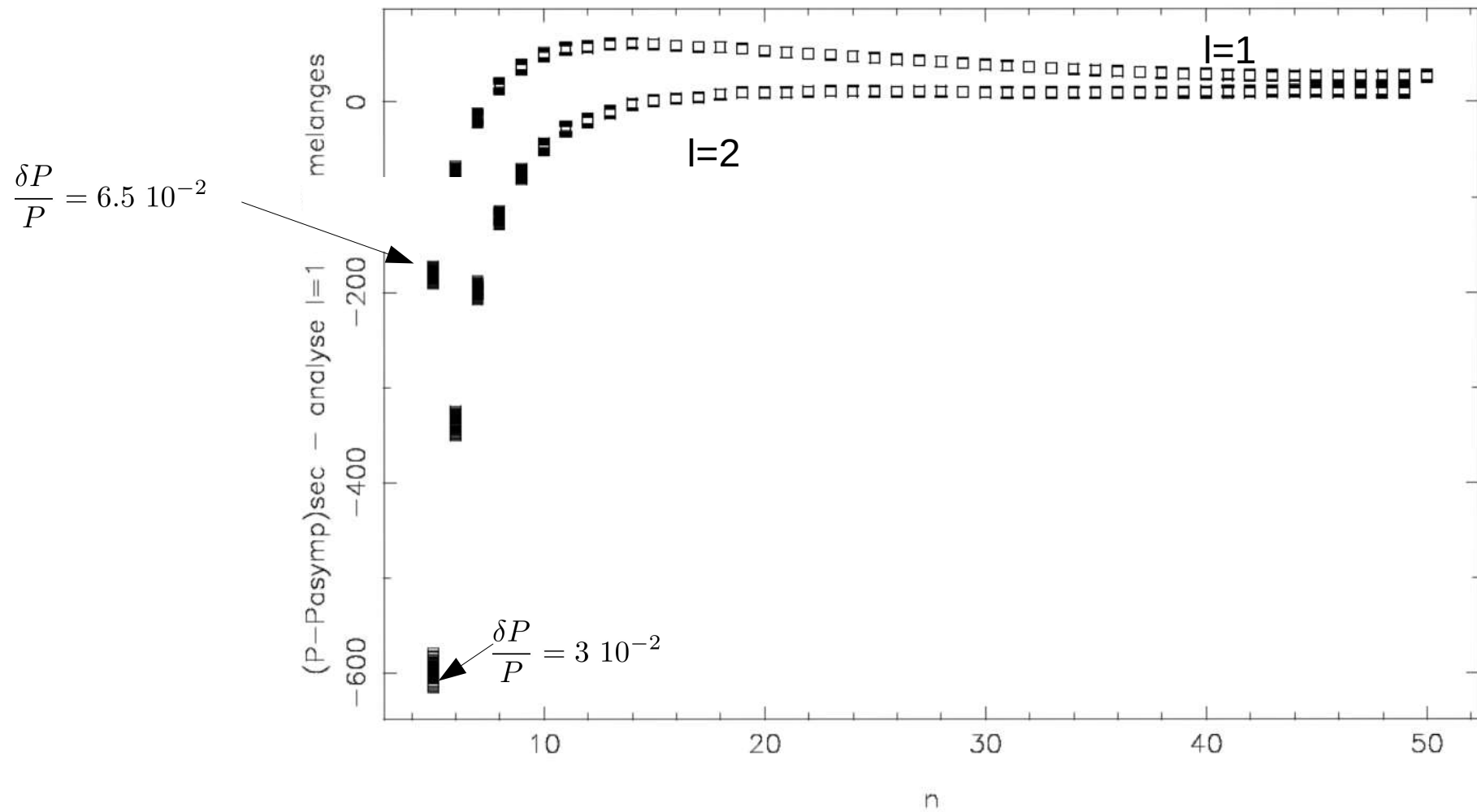
Fossat & Schmider 2018

$$\begin{array}{ll} \alpha_1 = 1100 \text{ s} & \\ \alpha_2 = 144 \text{ s} & \end{array} \quad \Rightarrow \quad \begin{array}{ll} V_1 \simeq -0.147 & (\sim 0.4) \\ V_2 \simeq 1.06 & (\sim 6) \end{array}$$

Comparision Numeric / Asymptotic



Comparision Numeric / Asymptotic



Asymptotic parameters from models

Two approaches :

(1) Fitting theoretical periods

=> model+oscillation code

(2) Directly from a model

=> compute V_1 , V_2 directly as a function of the model quantities

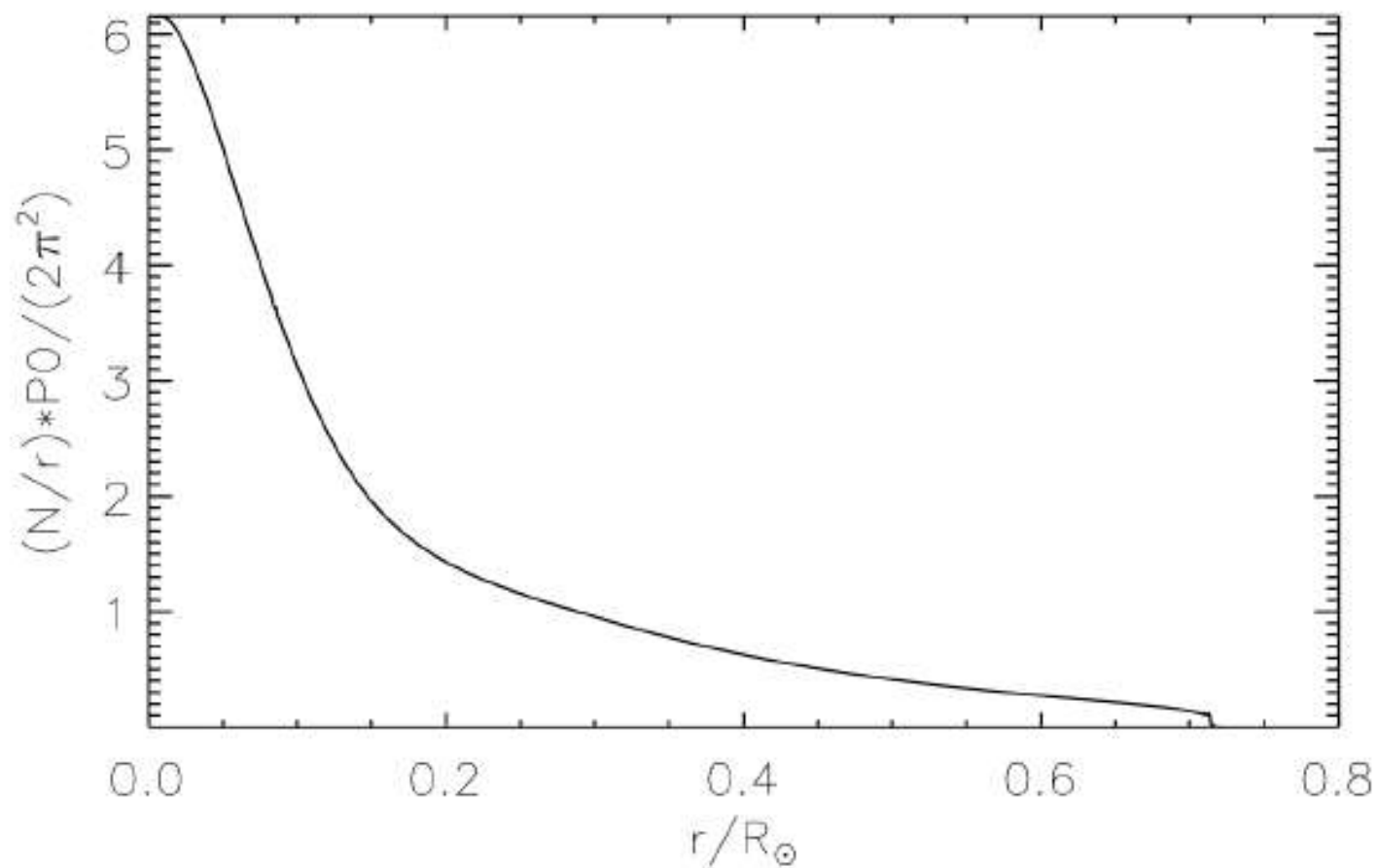
$$P_{n,l} = \frac{P_0}{\sqrt{l(l+1)}} \left[n + \frac{l}{2} - \frac{1}{4} - \theta \right] + \frac{P_0^2}{P_{n,l}} \frac{l(l+1)V_1 + V_2}{l(l+1)}$$

P_0

$$P_0 = 2\pi^2 \left(\int_0^R \frac{N}{r} dr \right)^{-1}$$

~35.355 mn

Lower bound of
previous histogram



$$V_1$$

$$V_1 = \frac{1}{P_0} \lim_{\epsilon \rightarrow 0} \left(\int_{\epsilon}^{r_c} \frac{dr}{Nr} - \frac{1}{N(\epsilon)} \right)$$

$$N(r) \underset{r \rightarrow 0}{\simeq} \alpha r$$

$$\int_{\epsilon}^{r_1} \frac{dr}{Nr} = \int_{\epsilon}^{r_1} \frac{dr}{\alpha r^2} = -\frac{1}{\alpha r_1} + \frac{1}{\alpha \epsilon} = -\frac{1}{N(r_1)} + \frac{1}{N(\epsilon)}$$

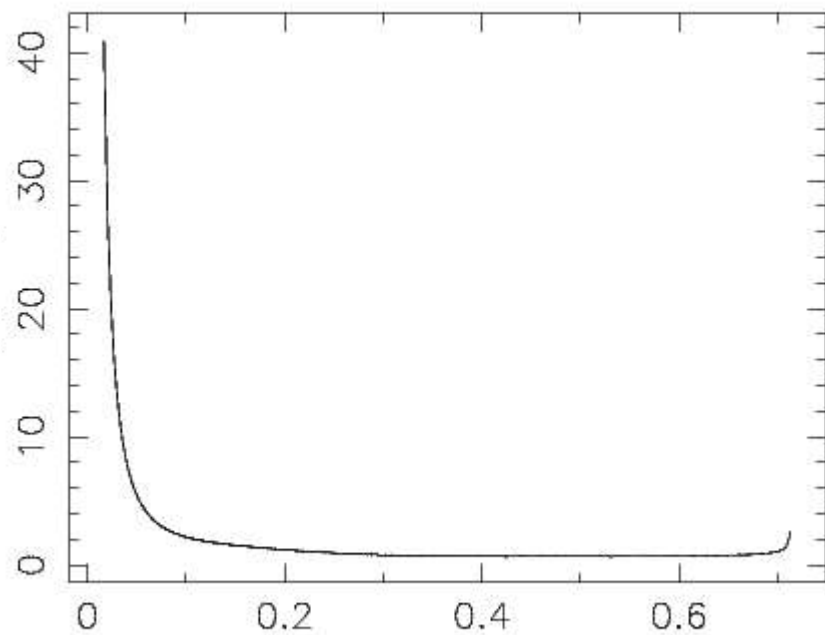
$$V_1 = \frac{1}{P_0} \left(\int_{r_1}^{r_c} \frac{dr}{Nr} - \frac{1}{N(r_1)} \right)$$

r_1 : first point of the model

V_1

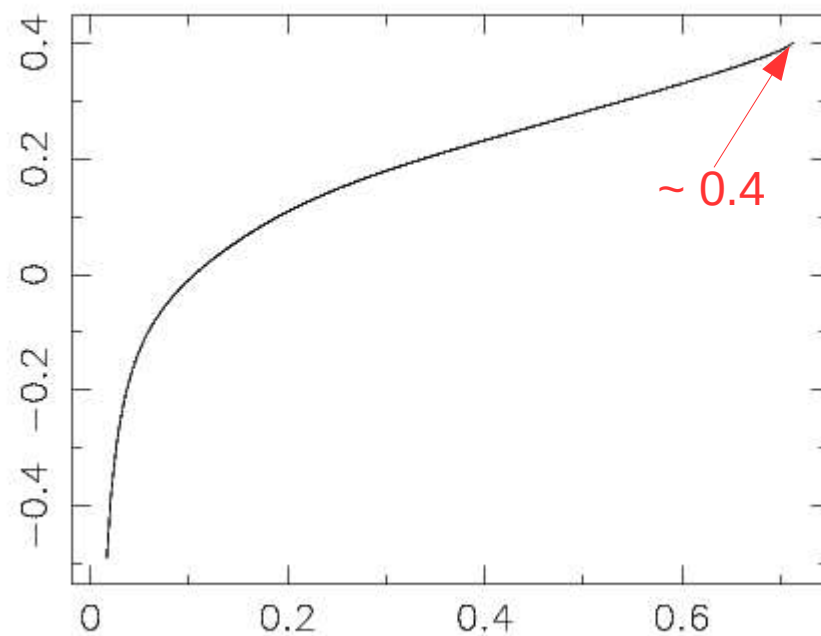
$$V_1 = \frac{1}{P_0} \left(\int_{r_1}^{r_c} \frac{dr}{Nr} - \frac{1}{N(r_1)} \right)$$

$1/(Nr)$



r/R

construction $V1 = [\text{int}_{r_{\min}}^{r_c} (dx/(x N)) - 1/N(r_{\min})] / P_0$



r/R

$$V_2$$

$$V_2 = \frac{1}{P_0} \lim_{\epsilon \rightarrow 0} \left(\int_0^{r_c - \epsilon} \left[\frac{1}{4} \left(\frac{\rho'}{\rho} \right)^2 - \frac{1}{2} \left(\frac{\rho'}{\rho} \right)' + A' - \frac{1}{16} \left(\frac{\varphi'}{\varphi} \right)^2 + \frac{1}{4} \left(\frac{\varphi'}{\varphi} \right)' \right] \frac{dr}{\varphi^{\frac{1}{2}}} + \frac{5}{24} \frac{1}{u(r_c - \epsilon)} \right) \quad (1)$$

$$\varphi = \frac{N^2}{r^2}; \quad A = \frac{\rho'}{\rho} - \frac{p'}{\Gamma p} = -\frac{N^2}{g}; \quad u(r_c - \epsilon) = \frac{3}{2} \int_{r_c - \epsilon}^{r_c} \varphi^{\frac{1}{2}} dr \quad (2)$$

$$V_2 = V_2^1 + V_2^2 \quad (3)$$

$$V_2^1 = \frac{1}{P_0} \left(\int_0^{r_c} \frac{r dr}{N} \left[\frac{1}{4} \left(\frac{\rho'}{\rho} \right)^2 - \frac{1}{2} \left(\frac{\rho'}{\rho} \right)' + A' \right] \right) \quad (4)$$

$$V_2^2 = \frac{1}{P_0} \lim_{\epsilon \rightarrow 0} \left(\int_0^{r_c - \epsilon} \left[-\frac{1}{16} \left(\frac{\varphi'}{\varphi} \right)^2 + \frac{1}{4} \left(\frac{\varphi'}{\varphi} \right)' \right] \frac{dr}{\varphi^{\frac{1}{2}}} + \frac{5}{24} \frac{1}{u(r_c - \epsilon)} \right) \quad (5)$$

$$V_2^2$$

$$V_2^2 = \frac{1}{P_0} \lim_{\epsilon \rightarrow 0} \left(\int_0^{r_c - \epsilon} \left[-\frac{1}{16} \left(\frac{\varphi'}{\varphi} \right)^2 + \frac{1}{4} \left(\frac{\varphi'}{\varphi} \right)' \right] \frac{dr}{\varphi^{\frac{1}{2}}} + \frac{5}{24} \frac{1}{u(r_c - \epsilon)} \right) \quad (5)$$

$$\frac{1}{4} \int_0^{r_c - \epsilon} \left(\frac{\varphi'}{\varphi} \right)' \varphi^{-\frac{1}{2}} dr = \frac{1}{4} \left[\frac{\varphi'}{\varphi^{\frac{3}{2}}} \right]_0^{r_c - \epsilon} + \frac{1}{8} \int_0^{r_c - \epsilon} \left(\frac{\varphi'}{\varphi} \right)^2 \frac{dr}{\varphi^{\frac{1}{2}}} \quad (6)$$

$$\left[\frac{\varphi'}{\varphi^{\frac{3}{2}}} \right]_{r=0} = 0 \quad (7)$$

$$V_2^2 = \frac{1}{P_0} \lim_{\epsilon \rightarrow 0} \left(\frac{1}{16} \int_0^{r_c - \epsilon} \left(\frac{\varphi'}{\varphi} \right)^2 \frac{dr}{\varphi^{\frac{1}{2}}} + \frac{1}{4} \left(\frac{\varphi'}{\varphi^{\frac{3}{2}}} \right)_{r_c - \epsilon} + \frac{5}{24} \frac{1}{u(r_c - \epsilon)} \right) \quad (8)$$

$$\boxed{\varphi(r) \underset{r \rightarrow r_c}{\simeq} \varphi_1(r_c - r); \quad \varphi' = -\varphi_1; \quad \frac{\varphi'}{\varphi} = -\frac{1}{r_c - r}} \quad (9)$$

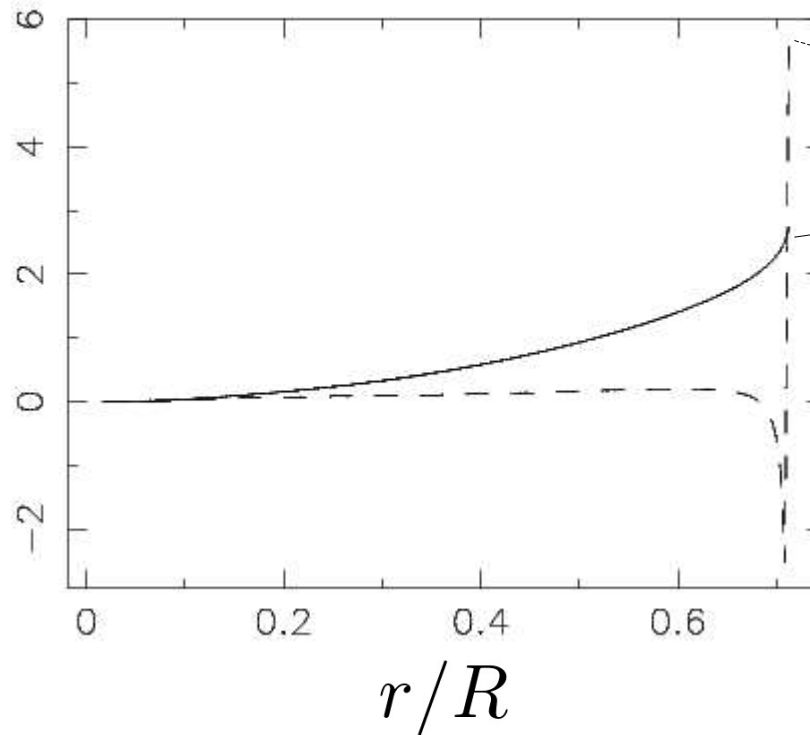
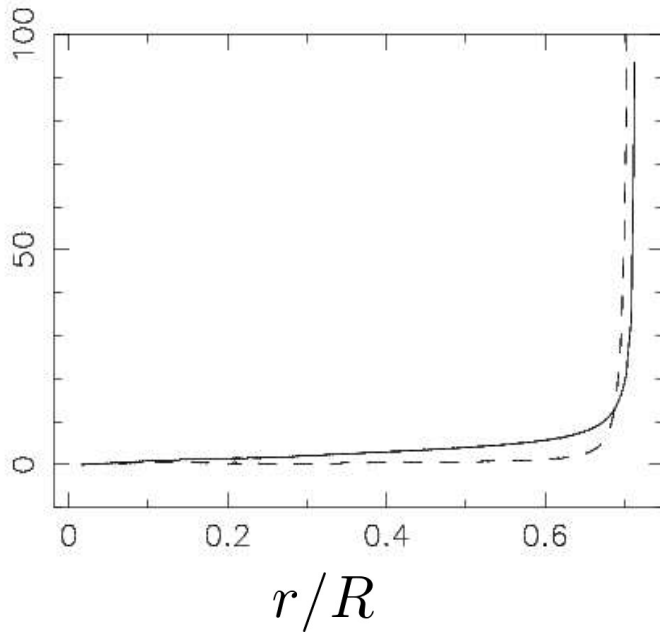
$$V_2^2 = \frac{1}{P_0} \left(\frac{1}{16} \int_0^{r_{\max}} \frac{\varphi'^2}{\varphi^{\frac{5}{2}}} dr - \frac{1}{24\varphi(r_{\max})(r_c - r_{\max})} \right)$$

$r_{\max}$: first point below the convective zone

V₂

$$V_2^1 = \frac{1}{P_0} \left(\int_0^{r_c} \frac{r dr}{N} \left[\frac{1}{4} \left(\frac{\rho'}{\rho} \right)^2 - \frac{1}{2} \left(\frac{\rho'}{\rho} \right)' + A' \right] \right)$$

somv2(full); somv2_2-1/24(dashed) su



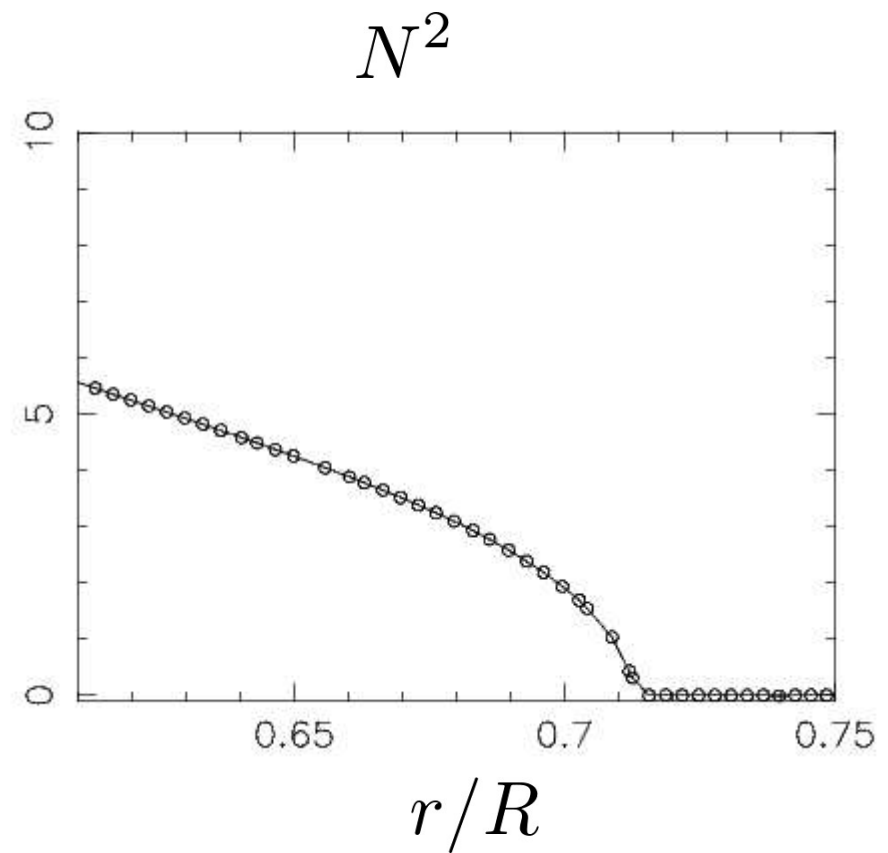
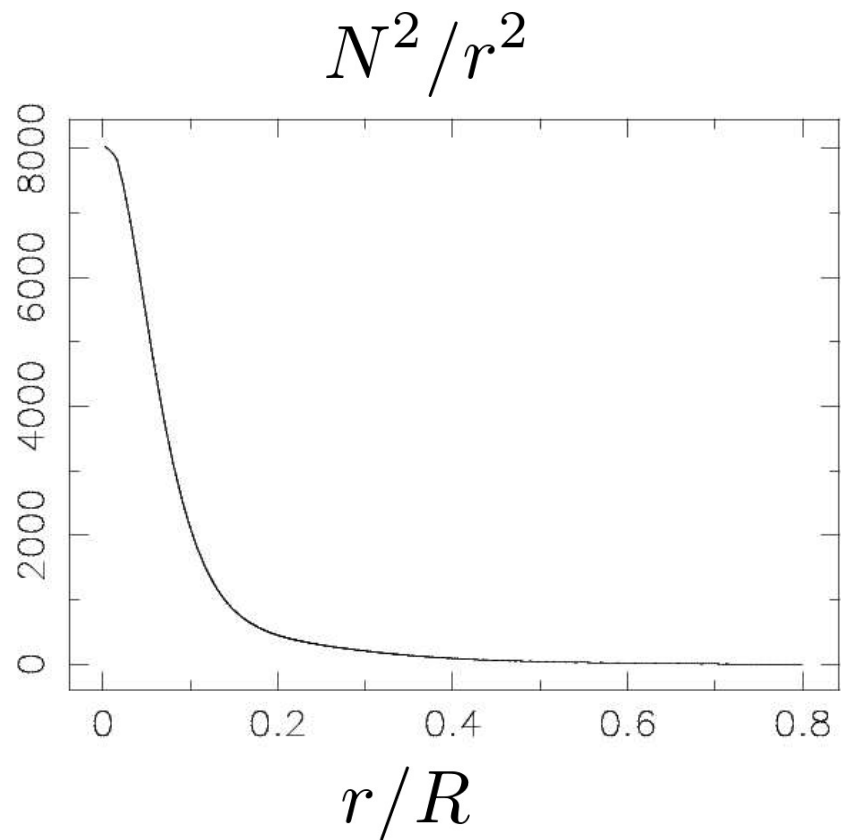
$$V_2^2 = \frac{5.57}{P_0} = 4.16$$

$$V_2^1 = \frac{2.73}{P_0} = 2.04$$

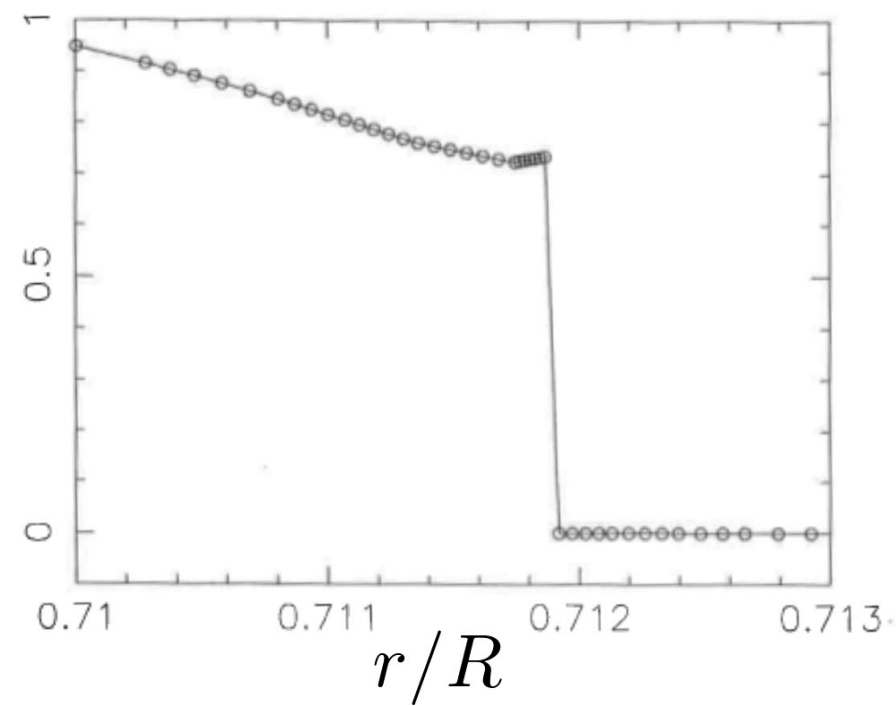
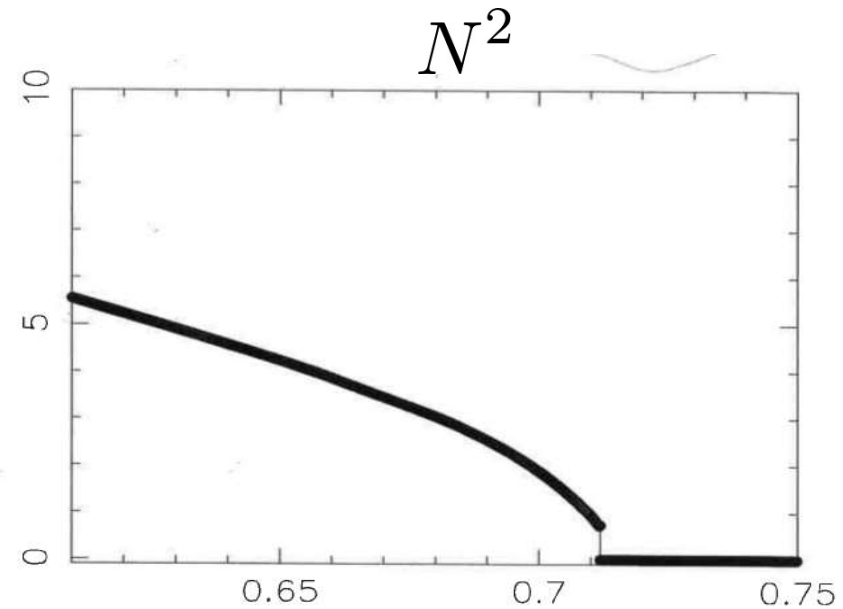
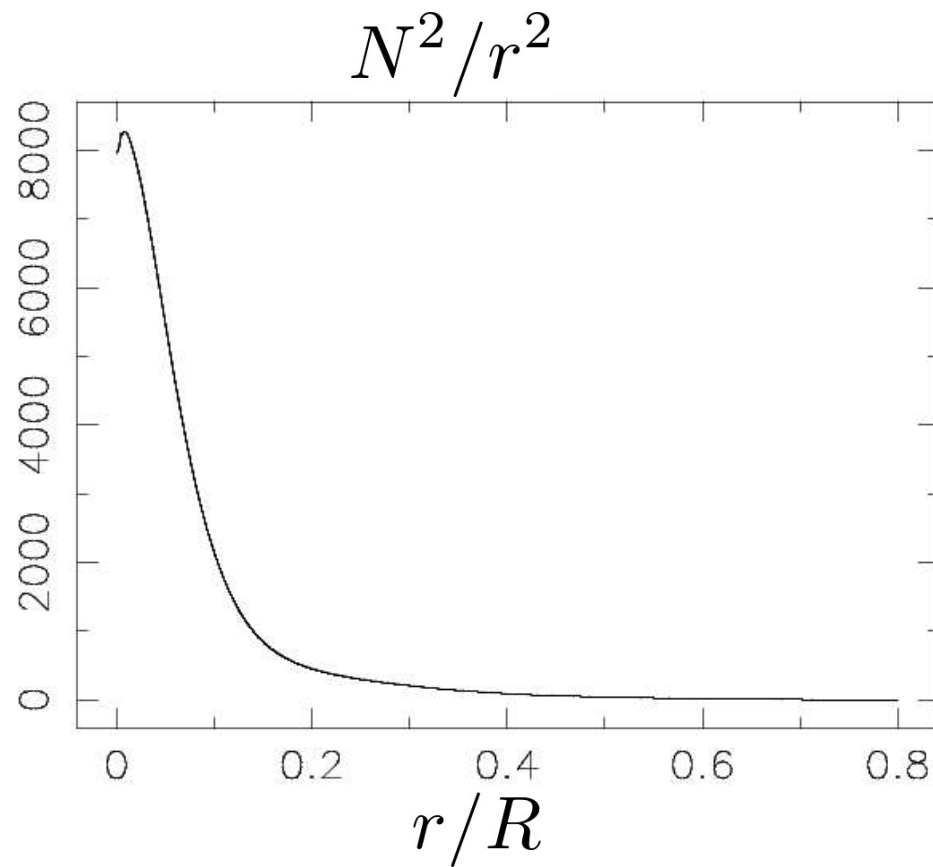
$$V_2 \simeq 6.2$$

$$V_2^2 = \frac{1}{P_0} \left(\frac{1}{16} \int_0^{r_{\max}} \frac{\varphi'^2}{\varphi^{\frac{5}{2}}} dr - \frac{1}{24\varphi(r_{\max})(r_c - r_{\max})} \right)$$

1000 points CESAM model



10 000 points CESAM model



Numerical errors (for a given model)

