

Weak Lensing Measurement of Stellar-to-Halo Mass Ratio with *DECaLS* Survey

Huanyuan Shan (SHAO)

Collaborators: *Eric Jullo, Anna Niemiec, Kaylan Burleigh,
Jean-Paul Kneib, Thomas Erben, Johan Comparat, Jean Coupon
and the DECaLS team*

Outline

- Background
- *DECaLS*: DECam Legacy Survey
- Preliminary results:
 - Stellar-to-Halo Mass Ratio (*SHMR*)
- Summary

SHMR: Stellar-to-Halo Mass Ratio



Aquarius Simulation

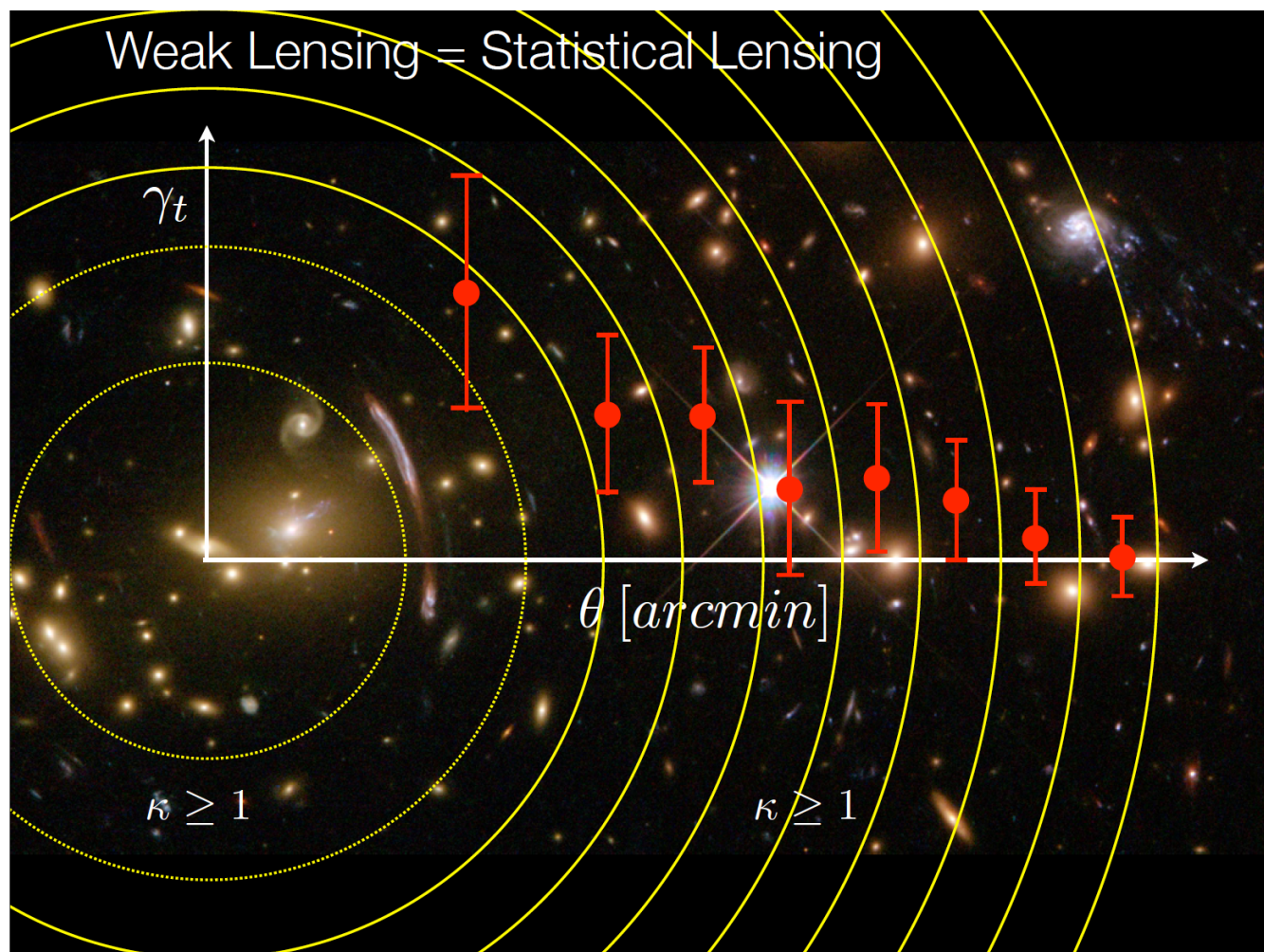
- ***SHMR***: $f_* = M_*/M_{\text{halo}}$
 - M_* : stellar mass
 - M_{halo} : halo mass

SHMR: Stellar-to-Halo Mass Ratio



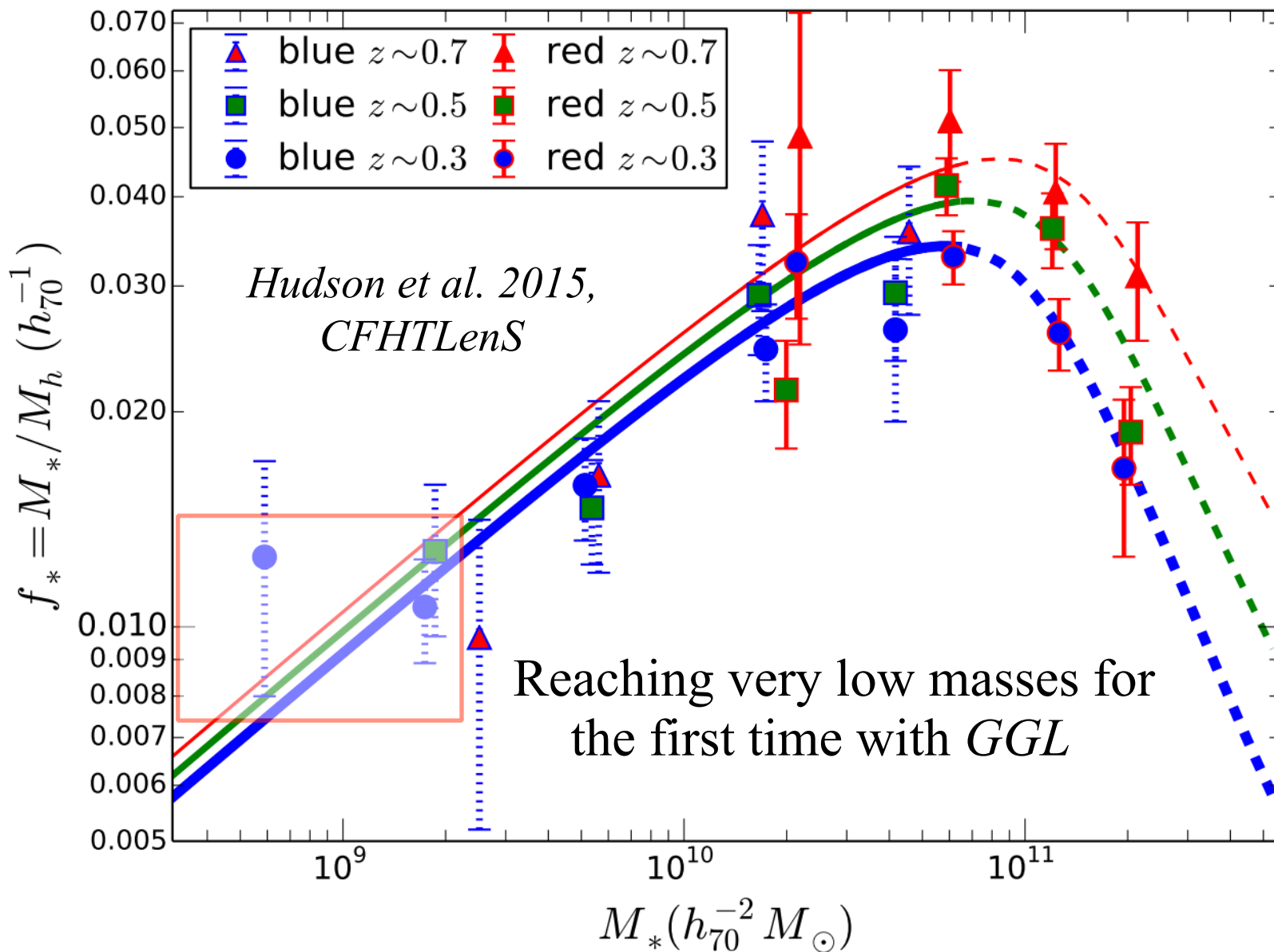
Aquarius Simulation

- Abundance matching
- Galaxy clustering
- Satellite kinematics
- Galaxy-galaxy lensing

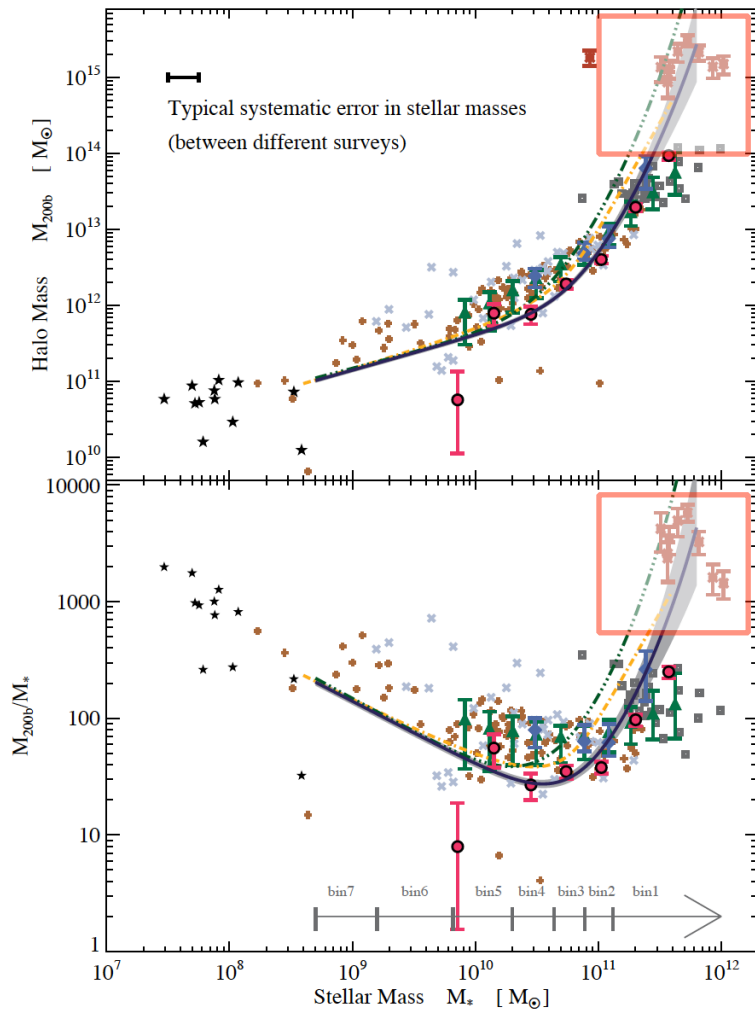


- Galaxy-galaxy lensing (**GGL**)
 - Measure DM halo mass directly
 - Reduce the noise (LSSs, shape noise, substructures, variation...)
 - Weak signal that is difficult to measure

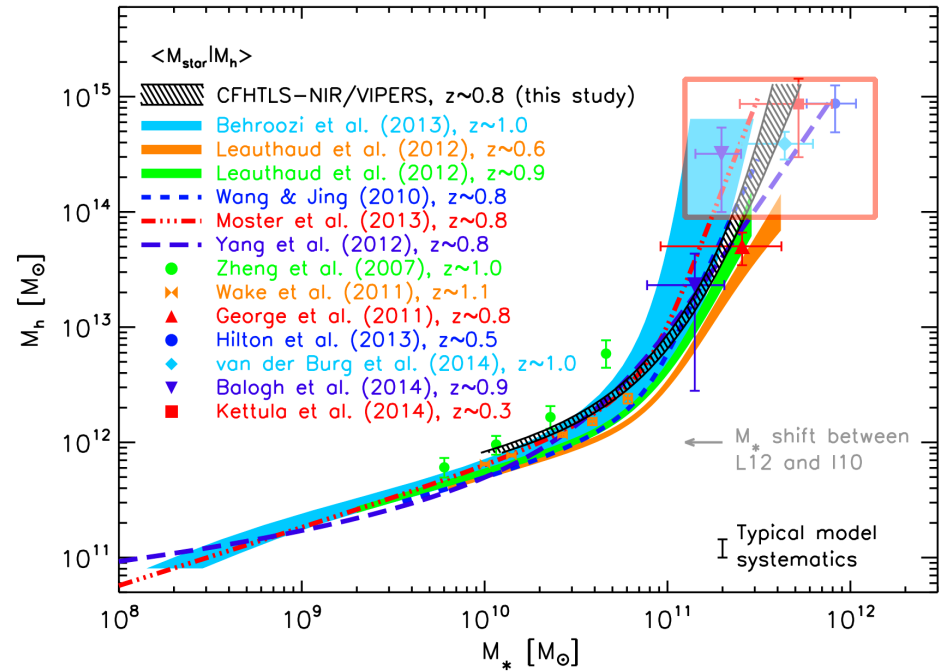
SHMR from GGL only



SHMR of high-mass halos



Coupon et al. 2015



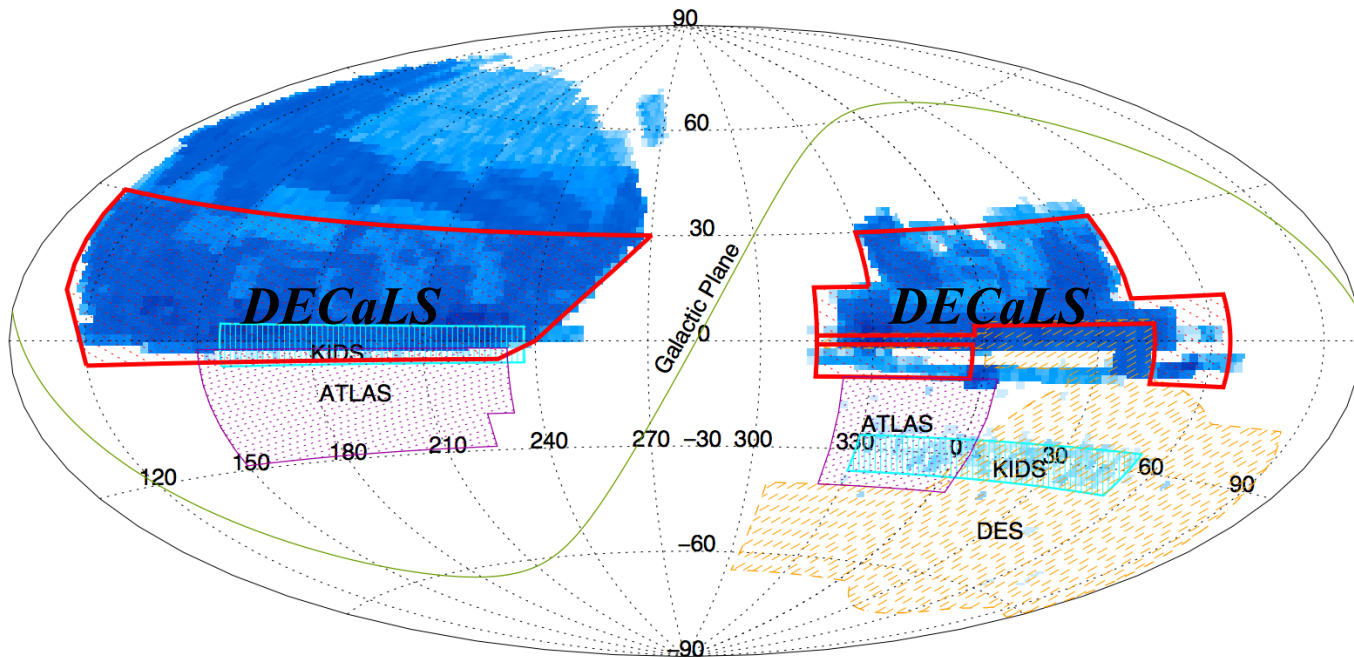
Leauthaud et al. 2012

SHMR of high-mass halos: **Large scatters**

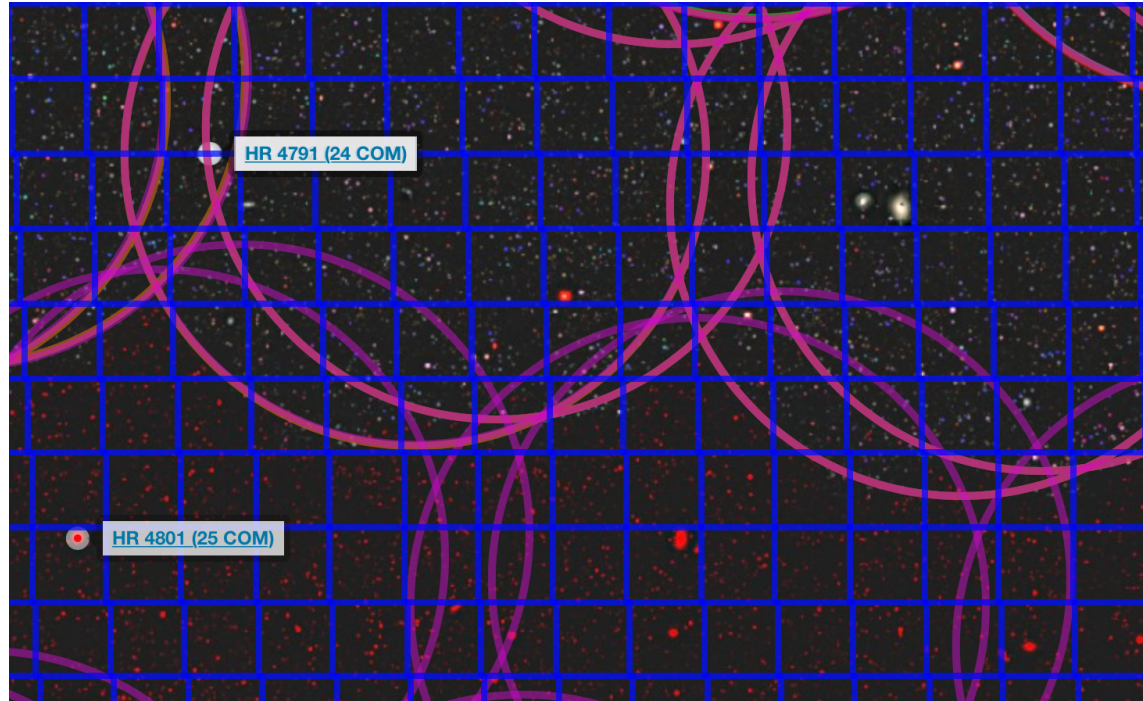
More constraints from *GGL*: **Larger WL surveys!**

DECaLS: DECam Legacy Survey

- Image *SDSS* footprint
 - Total area: $\sim 6700 \text{ deg}^2$
 - 3 bands: *g*, *r*, *z*
 - $\sim 1\text{-}2$ mag deeper than *SDSS* ($g \sim 24.7$, $r \sim 23.9$, $z \sim 23.0$)
 - Overlapping with the *BOSS/eBOSS/DESI* project



- Fov: 3 deg²
- 97554 bricks with the size 0.25*0.25 deg²
- Pixel size: 0.262''
- Mean seeing: 1.05''



Scientific Benefits

- GGL:
 - Massive spectroscopy of the *BOSS/eBOSS/DESI* project
 - Different samples: *LOWZ/CMASS, LRG, ELG, QSO*

Shape Measurement

- Data reduction: *Tractor* (Lang & Hogg)
- PSF modelling using *PSFEx* on each individual exposure
- Tracter shape measurement: PSF, SIMP, De Vaucouleur, exponential models are adjusted to every SExtractor source
 - De Vaucouleur and exponential profiles are decomposed in Mixture of Gaussians (MoG)
 - Pixelized PSF are convolved in Fourier Space

Morphological classification

Number of Sources	Type
478,918,959	Objects in a Primary brick
271,437,526	PSF
121,505,252	SIMP
63,568,420	EXP
20,141,591	DEV
2,266,170	COMP

Shape calibration

We bin the galaxies in mag_z and r_g . For each bin we compute

$$\chi^2 = \sum_i \frac{1}{\sigma_\gamma^2} \left(\bar{\gamma}_i^{\text{obs}} - [1 + m(\nu_{\text{SN}}, r)] \bar{\gamma}_i^{\text{true}} - c(\nu_{\text{SN}}, r) \right)^2$$

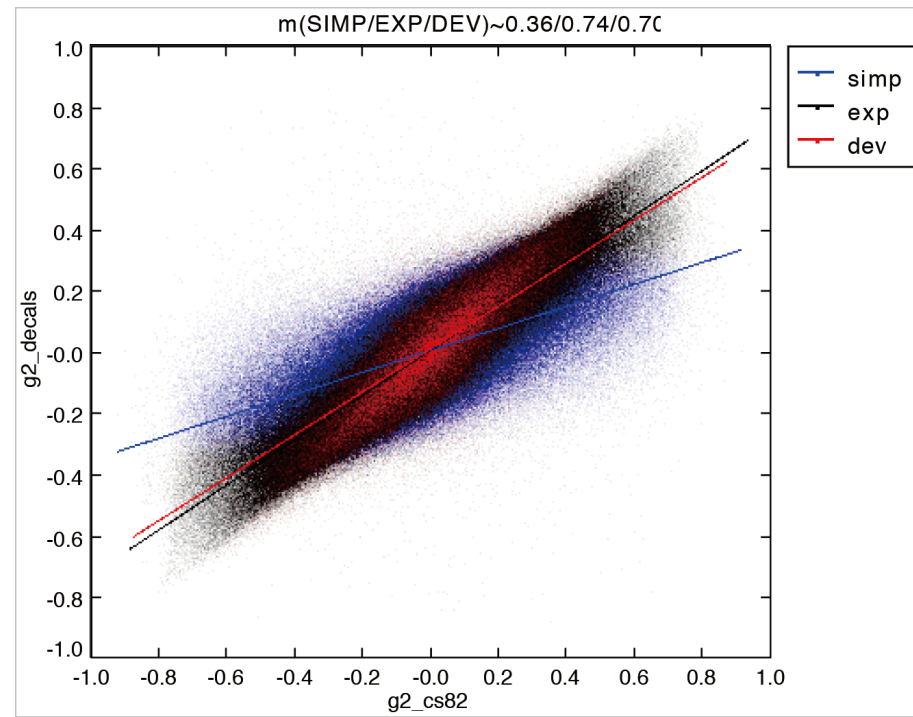
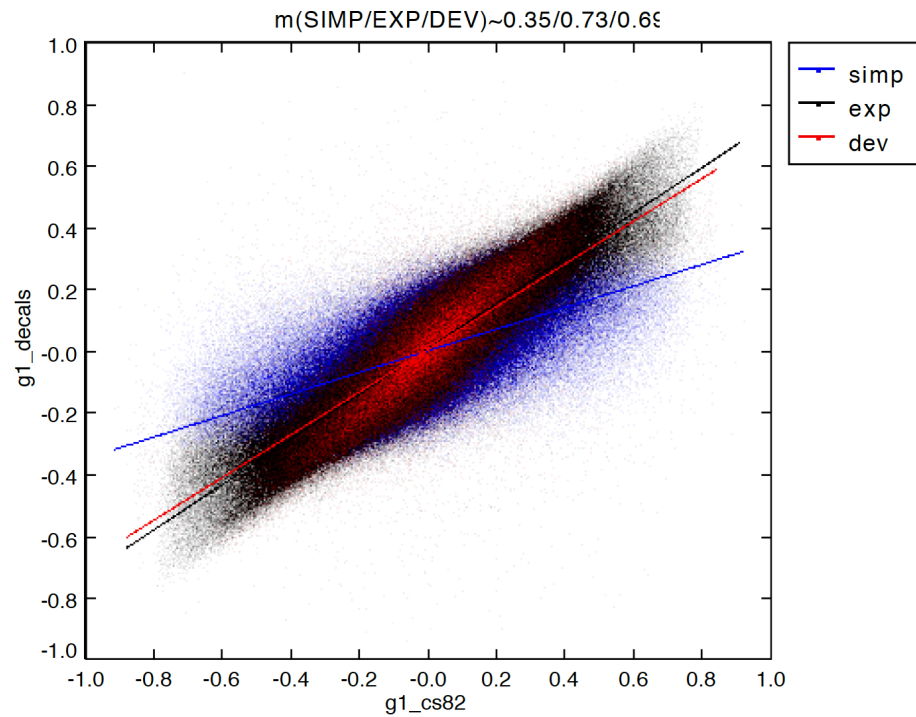
and adjust the functions $(1+m)$ and c as $\epsilon = (1 + m)\epsilon^{\text{true}} + c$

$$1 + m(\text{mag}_z, r_g) = a_0 \exp(-a_1 r_g \text{mag}_z) / \log_{10}(\text{mag}_z)$$

This is done for SIMP, EXP, DEV and COMP types

- Image simulation
- CS82 shear-calibrated data as the true values

- **m correction:** 10 magnitude bins & 10 galaxy size bins



- **c1, c2 correction** ($c=b_0+b_1*r_g+b_2*r_g^2$)
- No obvious trends above a few 10^{-4} as a function of magnitude and half light radius

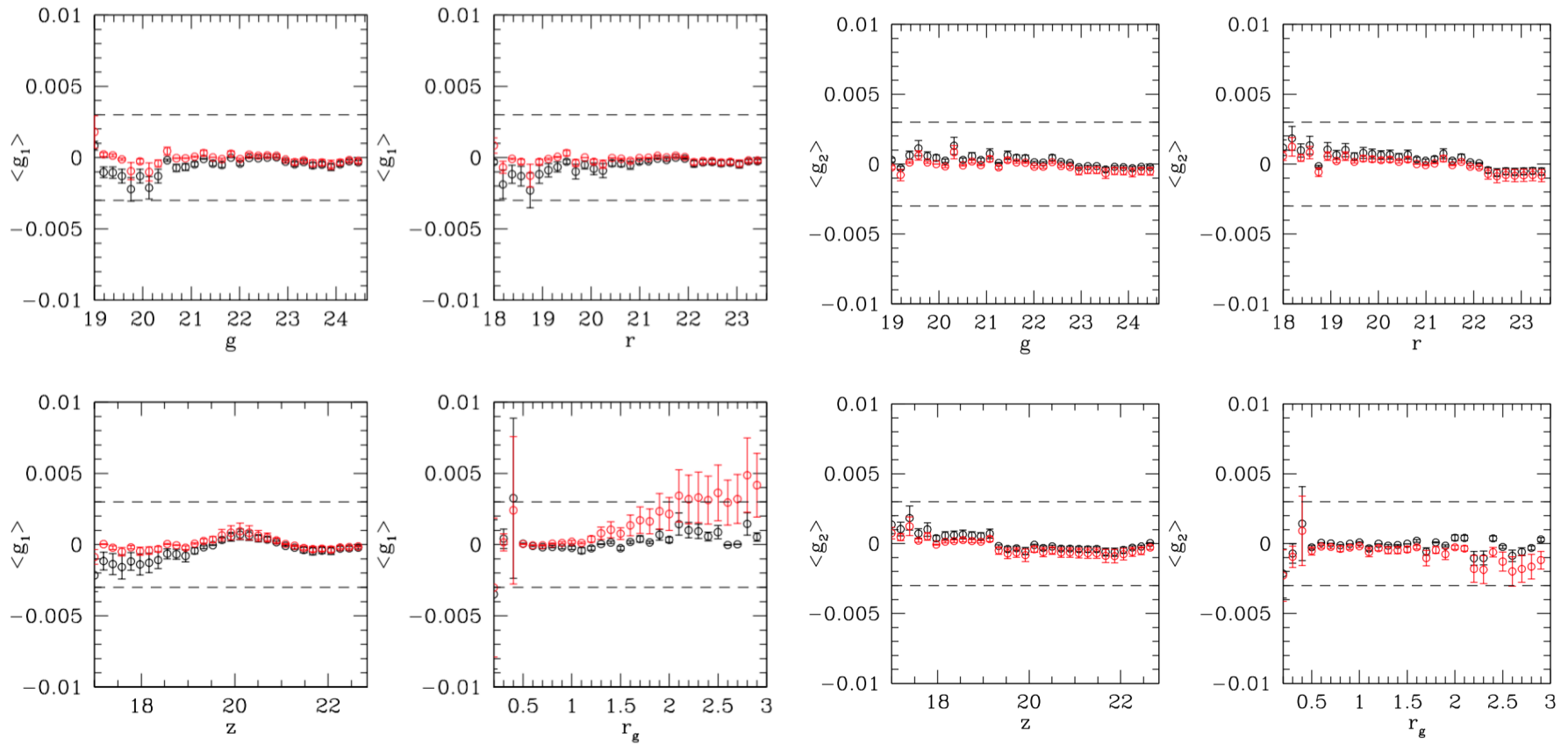
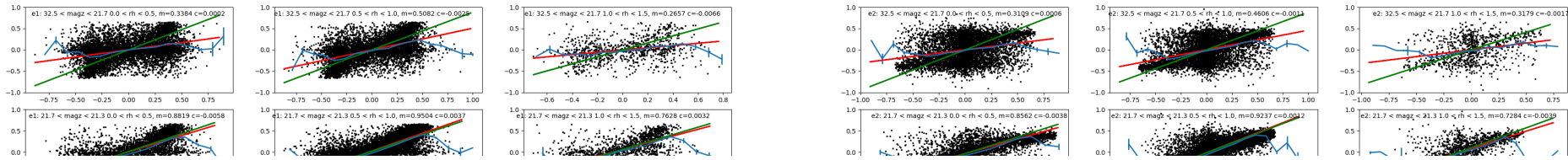
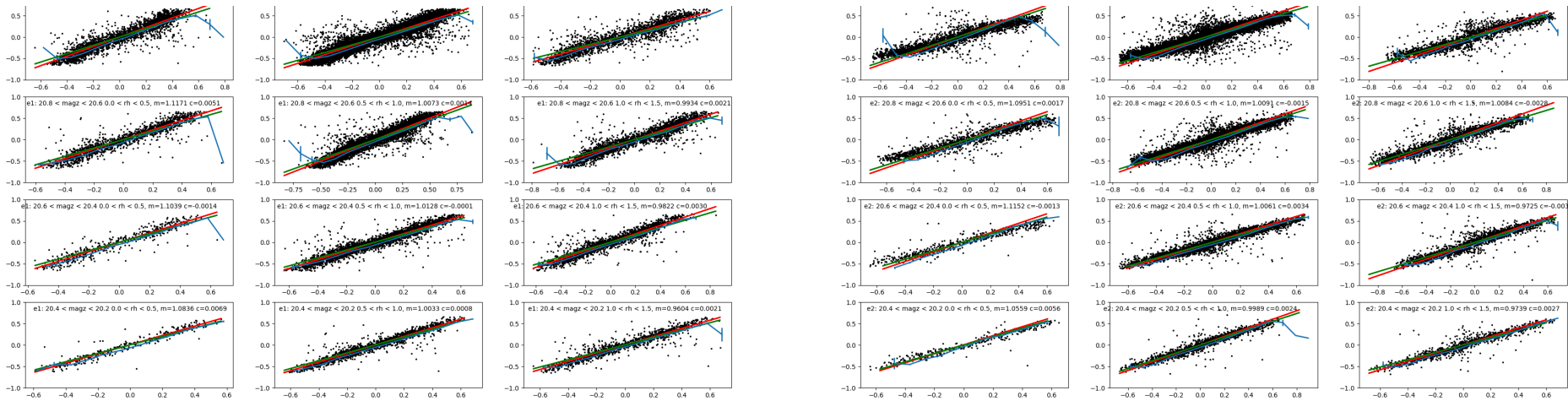


Image simulation:

- galaxy model+PSF+observation conditions
- DECaLS Legacy pipeline on simulated image

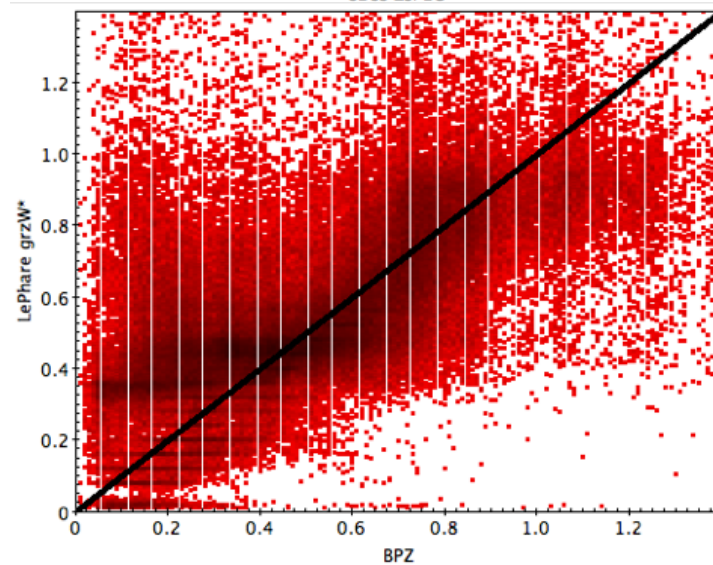
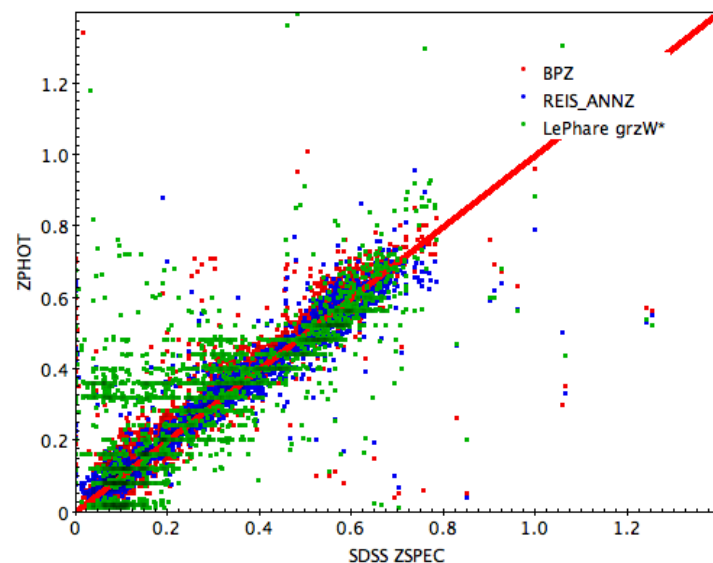


Similar m & c as comparison method



Photometric redshift

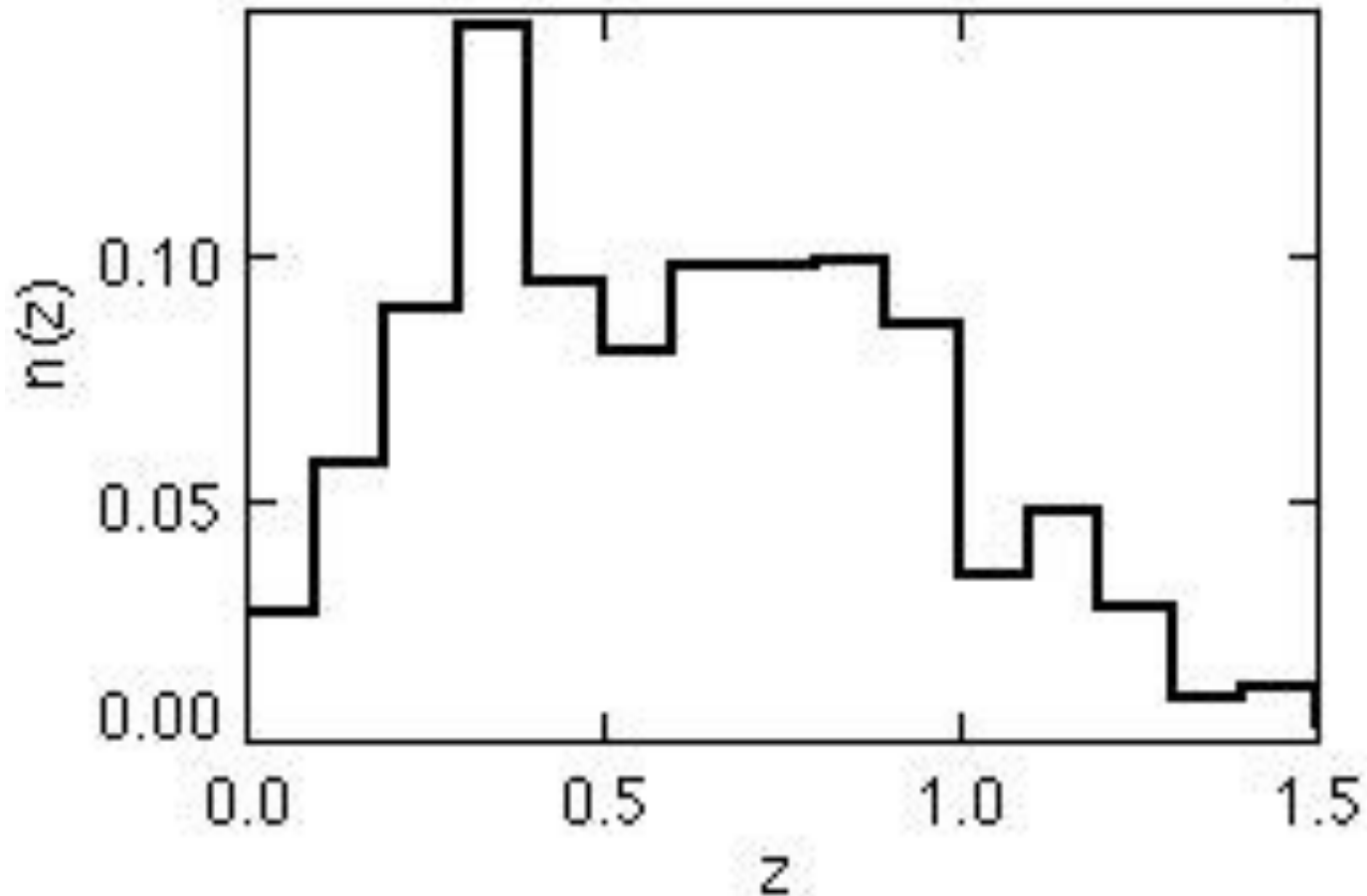
- We use the *GAZPAR* webservice provided by LAM to compute the *z*-phot of DECaLS galaxies using *grzW* bands
 - We compare to CS82 *z*phots
- No particular systematic error, but a *large scatter*



- **Galaxy selection:**

$19.0 < \text{mag}_g < 24.65$ & $18.0 < \text{mag}_r < 23.61$ & $17.0 < \text{mag}_z < 21.0$ & $0.65 < r_g < 3.0$

- **$n(z)$ from *COSMOS***

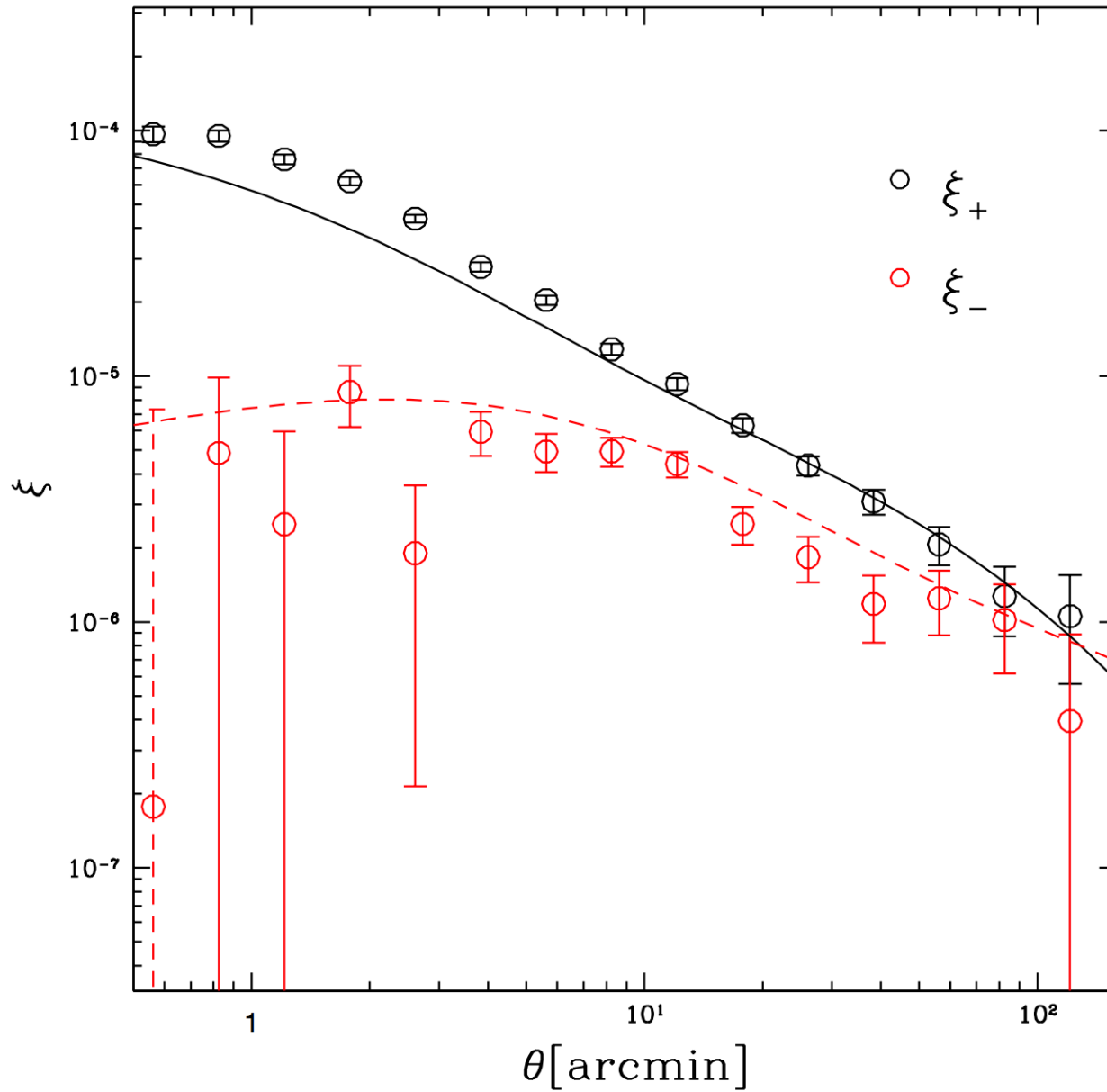


DECaLS: Lensing Status

- DECaLS DR3: 4300 deg² in *g*-band, 4600 deg² in *r*-band, 8100 deg² in *z*-band
- 4200 deg² have been observed in 3 bands

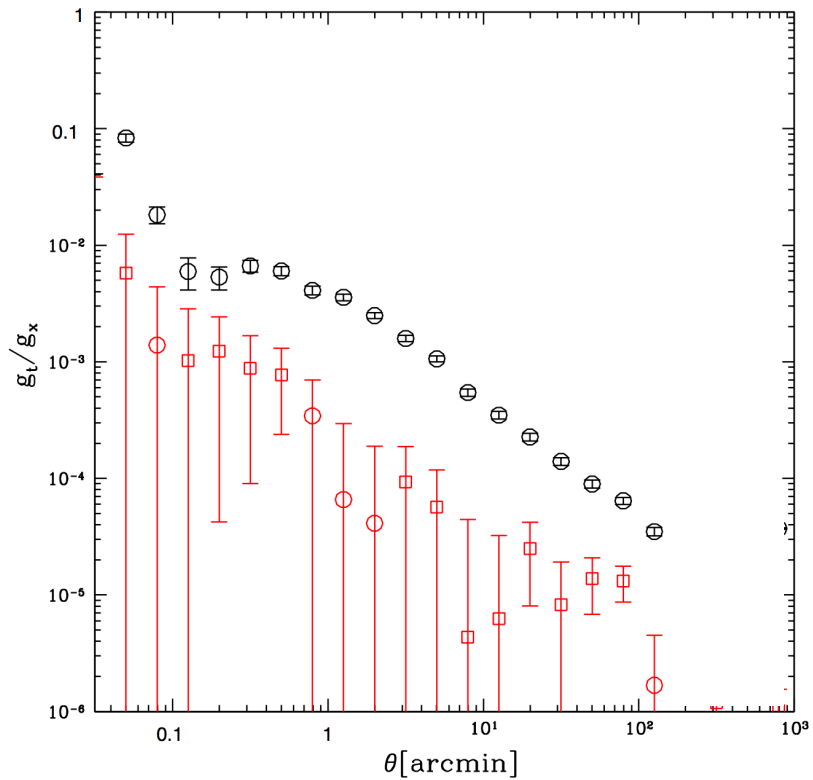
- About 56 million galaxies useful for lensing
- Lensing galaxy density about 4 arcmin⁻²

Shear correlation function

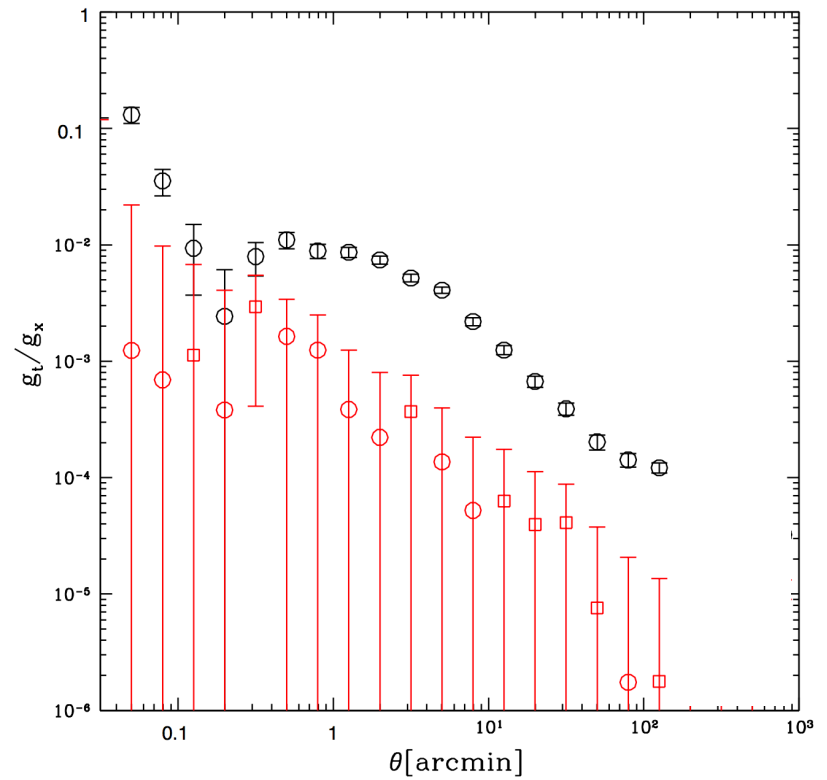


Tangential shear

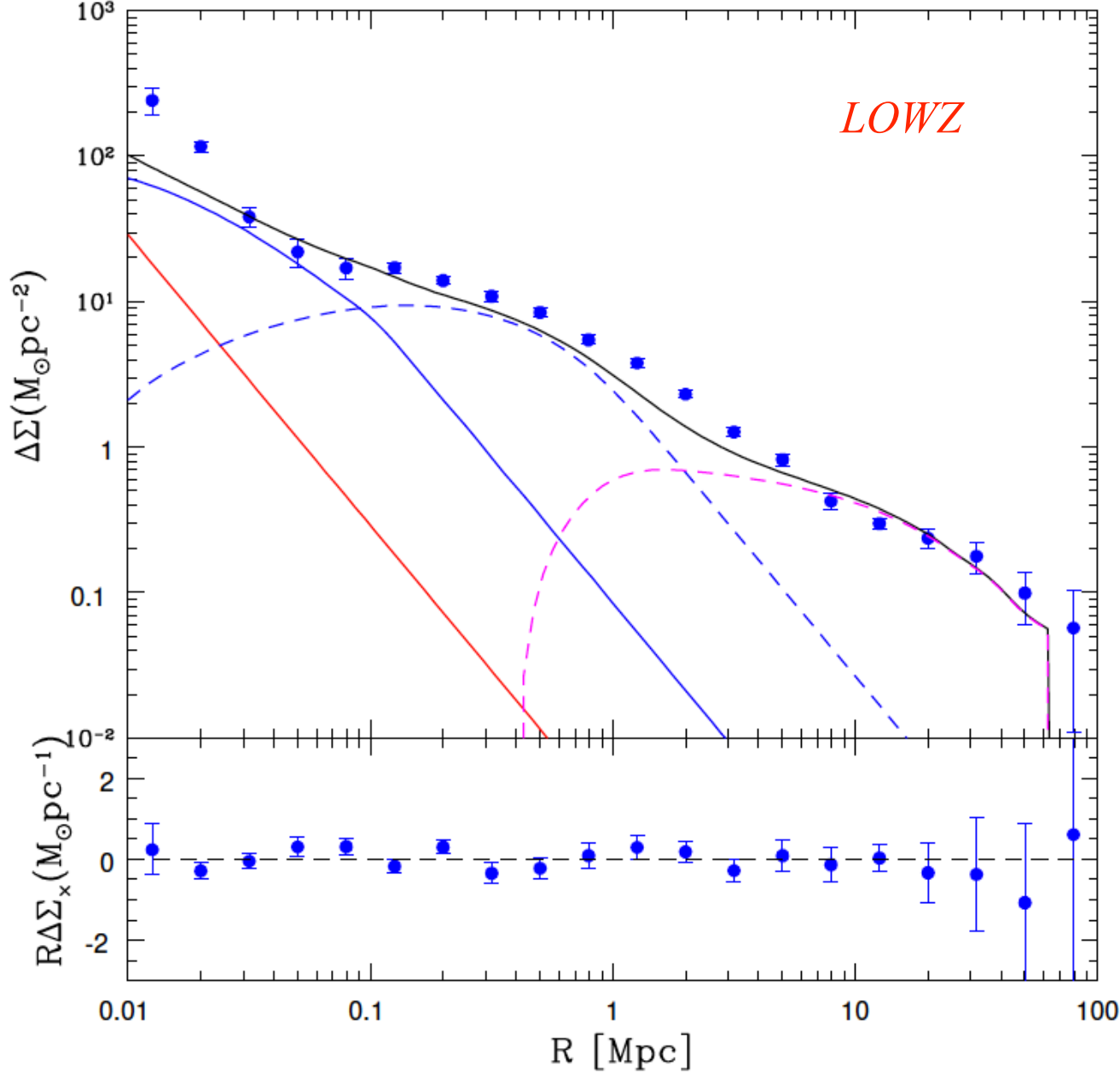
LOWZ galaxies (BOSS DR12)



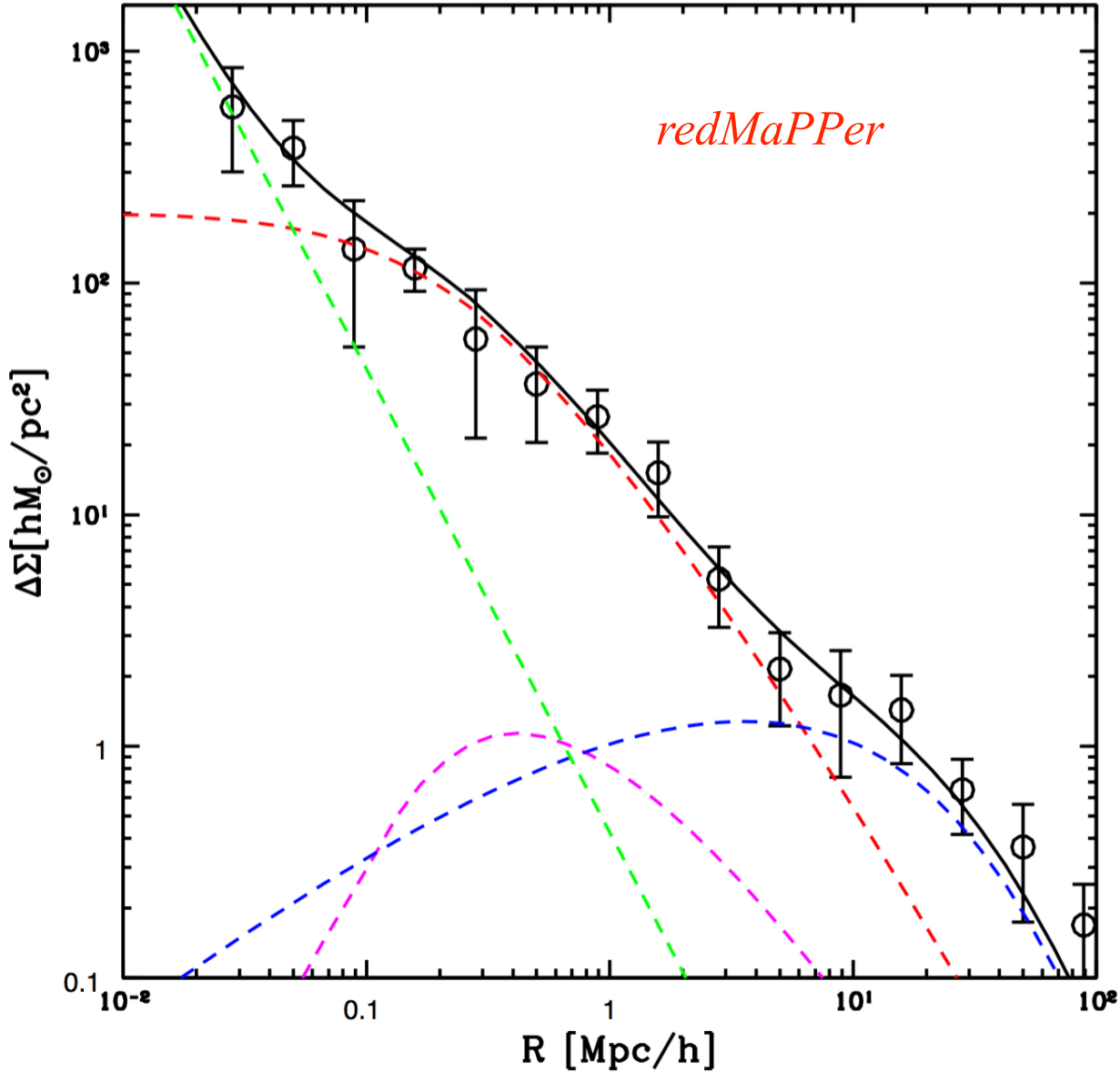
redMaPPer clusters
(*Rykoff et al. 2014*)



GGL: Stellar mass+NFW halo+satellite+2-halo

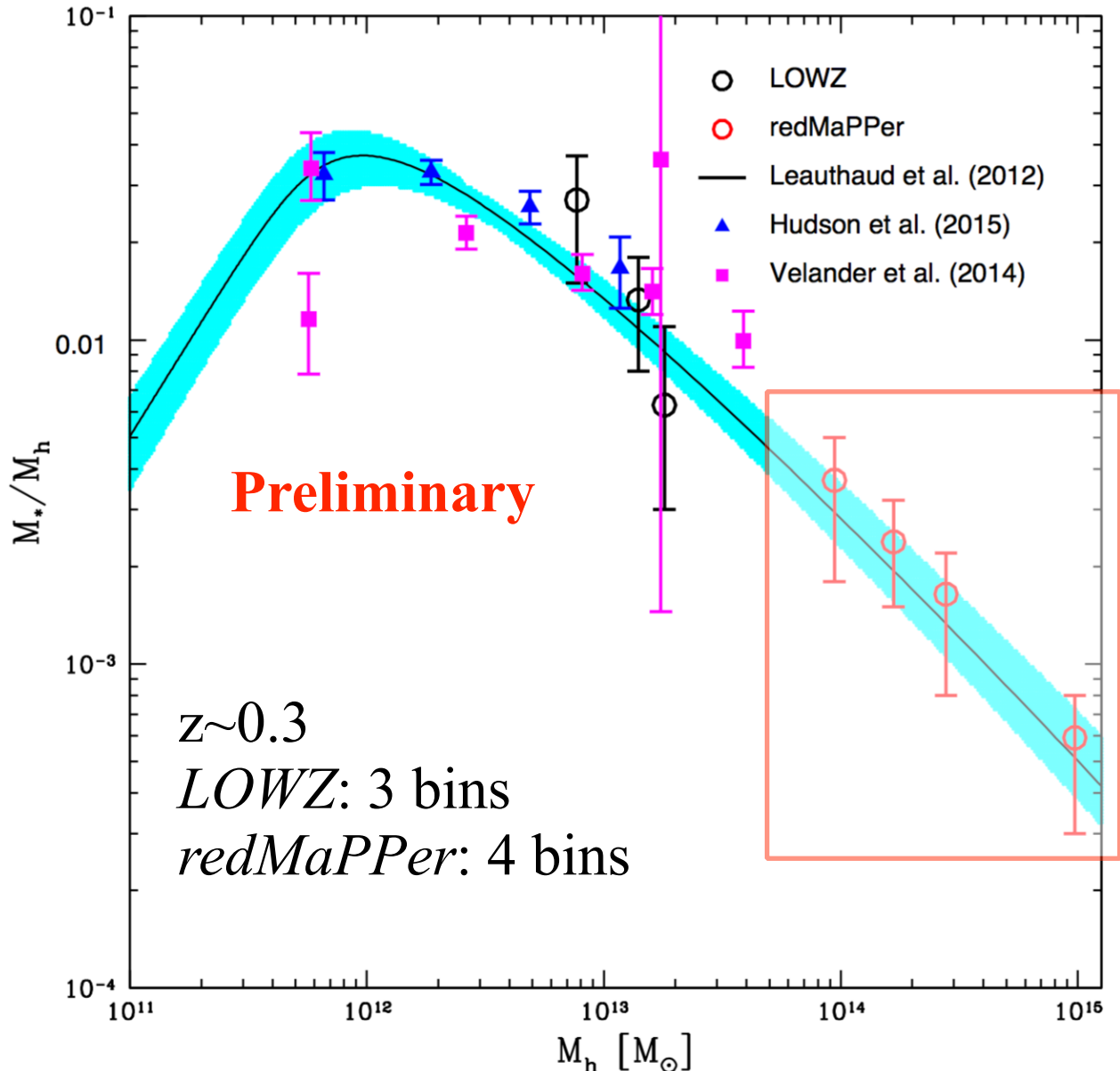


GGL: Stellar mass+NFW halo+Mis-centering+2-halo



SHMR of central galaxies

- Fixed stellar mass (*Maraston et al. 2013, BOSS*)



Summary

- *DECaLS*
 - Massive spectroscopy of the *BOSS/eBOSS/DESI* projects
- Measurement
 - *Tractor*: shear catalog
 - *COSMOS*: redshift distribution
- Preliminary results:
 - GGL signals of LOWZ/redMaPPer
 - *SHMR*